

1 **Description and assessment of neuston communities present in the**
2 **South Indian Ocean**

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4 By

5
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13 UNIVERSITY of the
14 WESTERN CAPE

19 **Description and assessment of neuston communities present in the South Indian Ocean**

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22 Western Cape.

23 **Abstract**

24 Animal communities living at the sea-surface are relatively poorly known, despite the fact
25 that this biotope is globally distributed, it acts as an important nursery area for many
26 commercially important taxa and supports significant amounts of anthropogenic waste. This
27 lack of knowledge applies particularly to the Indian Ocean. Here, I describe the structure of
28 non-copepod neuston assemblages across the Southern Indian Ocean Gyre using samples
29 collected between 26 June – 16 July 2015 from 35 stations along a transect line at 20° S.
30 Neuston was sampled using triplicated manta trawls at each station, while environmental
31 parameters were assessed using a CDT. All non-copepod neuston were identified to
32 morphospecies, with the exception of siphonophores, which were identified to species.

33 At the time of the study, two anticyclonic eddies were present between 75 – 85° E and these
34 were characterised by warmer surface waters, deeper fluorescence maxima and elevated total
35 fluorescence, by comparison with surrounding areas. Salinity was lowest to the east.
36 Assemblages were dominated by Ostracods, though mysids, juvenile chaetognaths,
37 siphonophores and amphipods were also common: the majority of specimens collected were
38 facultative neuston. The greatest number of morphospecies were observed in Amphipoda and
39 Siphonophora. While facultative neuston were more common at night than by day, euneuston
40 showed little diel variation in abundance. Meroplanktonic neuston were more abundant at
41 longitudinal extremes and were uncommon in the centre of the area.

42 There was a significant spatial pattern to the structure of assemblages, with those in the centre
43 of the transect (between 75 - 85° E) being distinct from those to either the east or west.
44 Samples collected in the far east of the region were different from the balance of samples,
45 which tended to show some rotational symmetry around assemblages at the core. The key
46 environmental factors that were linked to assemblage structure were associated with
47 longitude – salinity, fluorescence and temperature at 200 m, emphasising the interaction
48 between Indonesian Throughflow Water, eddy-related process and settlement of

49 meroplankton. The study highlights the value of using morphospecies in studies of plankton
50 assemblages.

51

52

53 **Keywords:** Air-sea interphase, biogeography, biological communities, Indian Ocean,
54 neuston, plankton, Subtropical Gyre

55



56

57 **Declaration**

58 **I, Yasmeen Parker declare that:**

59 “Description and assessment of neuston communities present in the South Indian Ocean”, is
60 my own work, and that:

- 61 1. It has not been submitted for any degree or examination at any other university;
- 62 2. When data (neuston samples) were not collected by me, the sources were
63 acknowledged;
- 64 3. All the literature sources that I have used and quoted have been acknowledged in my
65 reference list.

66

67 Full name: Yasmeen Parker

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Signature:



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Date: 30 November 2021

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128

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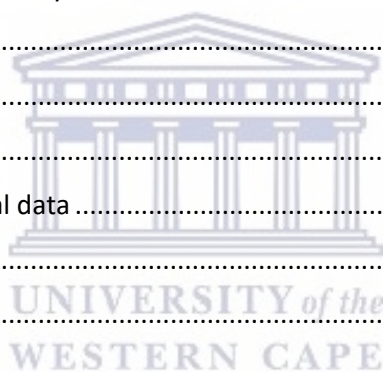
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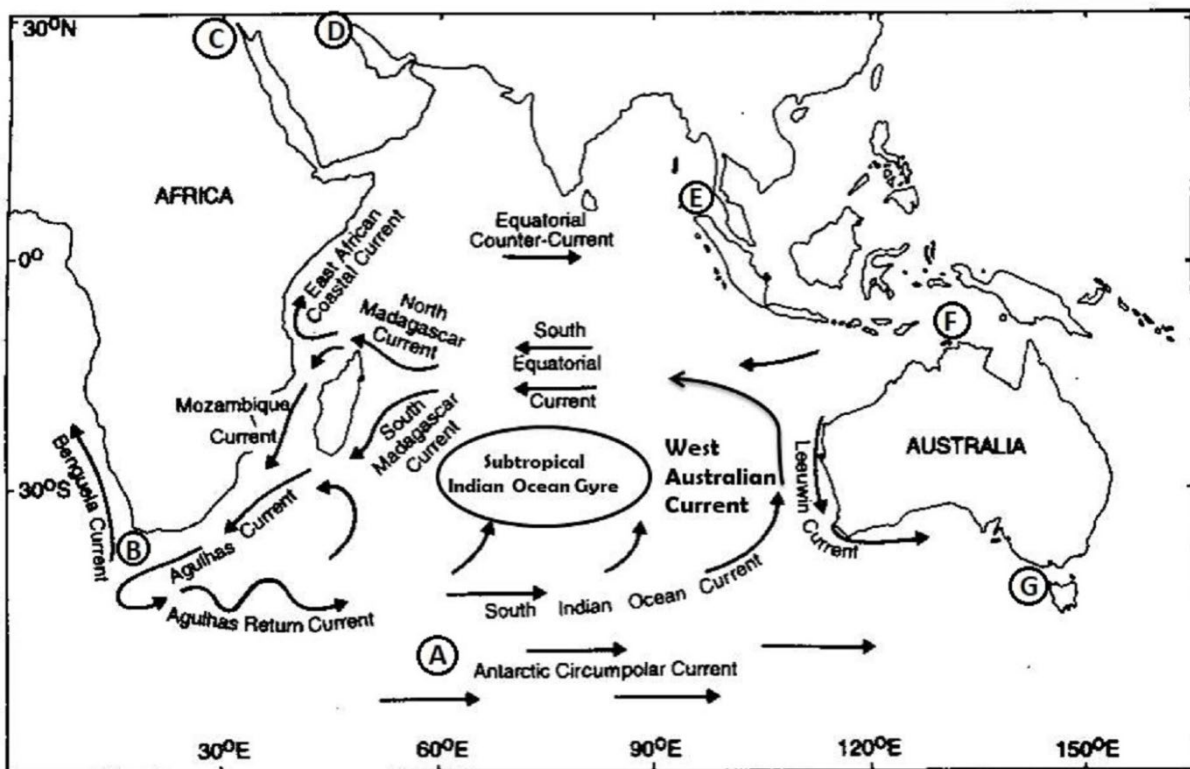
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157 **Introduction**

158

159 The Indian Ocean is the world's third- largest ocean, covering an area of over 68.5 million
 160 km², and is bounded by land masses on three sides. The Indian Ocean stretches 9 600 km
 161 from the Bay of Bengal in the North to Antarctica in the South, and 7 600 km from Africa to
 162 Australia (Demopoulos et al. 2003) encompassing seven geographical limits (Figure 1, A-G).
 163 The Indian Subtropical Gyre Province (ISSG) extends from the hydrochemical front to the
 164 Subtropical Convergence (10-15° S to approximately 30° S) (Longhurst 1998).



165

166 **Figure 1:** Map of the Subtropical Indian Ocean Gyre showing the currents and the
 167 geographical limits of the Indian Ocean. **A**, The northernmost limit of the Southern Ocean at
 168 a latitude of 60° South, **B**, the southernmost point of the African continent at Cape Agulhas, **C**,
 169 the Suez Canal, **D**, the northernmost regions of the Persian Gulf, **E**, the Strait of Malacca, **F**,
 170 from the Singapore Strait to the eastern boundary of the Timor Sea (this includes the Java,
 171 Flores and Banda Seas within the archipelagic waters of Indonesia) and, **G**, the west coast of
 172 Tasmania (not including the waters of Bass Strait)(Pearce 1993, Allan et al. 1995, O'Loughlin
 173 and Luke 2010)

174 The Indian Ocean region encompasses 36 key nations and supports more than 25%
175 (O'Loughlin and Luke 2010) of the world's ever-increasing population (Hood et al. 2016),
176 while it's nine marine ecosystems house over 35 000 marine species (Colgan 2017). A
177 number of sub regions within the Indian Ocean have previously been recognised as having
178 very high levels of biodiversity (Colgan 2017). This biodiversity is threatened by a number of
179 factors including global warming and ocean acidification (Hood et al. 2016), as well as more
180 immediate pressures associated with the increasing human populations in the bordering states
181 (Hood et al. 2016). It is due to these stressors that understanding and predicting the changes
182 in the Indian Ocean have become imperative within recent years (Hood et al. 2016).

183 The Indian Ocean was identified as one of the least studied ocean basins by the Scientific
184 Committee on Ocean Research (SCOR) and the Intergovernmental Oceanographic
185 Commission (IOC) in the middle part of the twentieth century (Hood et al. 2016). As a
186 consequence, these bodies motivated for the International Indian Ocean Expedition (IIOE),
187 which ran from 1960 until 1965 (Hood et al. 2016). It was argued that human societies in this
188 region would benefit from studies aimed at understanding the annual monsoon cycle and its
189 wider effects on the Indian Ocean fisheries and weather (Hood et al. 2016).

190 The outputs from the IIOE included advancements on understanding monsoon variability and
191 dynamics, the emergence of the first description of complex surface ocean circulation in
192 response to monsoon forcing and a more detailed understanding of the complex topography
193 of the Indian Ocean Basin. Other outputs include a comprehensive hydrographic atlas of the
194 Indian Ocean basin (Wyrcki 1971, Schott and McCreary 2001) and the first survey of
195 monsoon circulation in the Somali Current (Swallow and Bruce 1996, Schott and McCreary
196 2001). A number of studies have been conducted on the Somali, Agulhas and Leeuwin
197 Currents (Stramma and Lutjeharms 1997), and these have served to improve our
198 understanding of the oceanography in the boundary regions of the Southern Indian Ocean.
199 However, our understanding of the dynamic processes such as the wind-driven, anticyclonic
200 circulation of the Southern Indian Ocean, which is significantly different to that of the South
201 Atlantic Ocean, remains weak (Peterson and Stramma 1991, Stramma and Lutjeharms 1997).

202 Since their first expedition to Somalia in 1975, the Norwegian research vessel *Dr Fridtjof*
203 *Nansen* has undertaken 40 surveys in the Western Indian Ocean, with the history of the
204 surveys being reviewed by Groeneveld and Koranteng (2017). The Western Indian Ocean is
205 dominated by two western boundary currents, the Agulhas to the south and the Somali to the

206 north (Groeneveld and Koranteng 2017). The Western Indian Ocean consists of dynamic
207 ocean currents and upwelling cells that regulate the climate and influence the weather
208 patterns, sea temperature, water chemistry, productivity, biodiversity and fisheries
209 (Groeneveld and Koranteng 2017). Mesoscale eddies also influence biological production by
210 interacting with the continental slope on the western side of the Mozambique Channel which
211 causes the upwelling of cooler, nutrient-rich, waters. (Ternon et al. 2014; Lamont et al. 2014;
212 Hugget and Kyewalyanga 2017).

213 Global winds driven by latitudinal pressure gradients, transfer energy to surface waters to
214 create shallow currents which, under the influence of the Coriolis force and through
215 interactions with continental land masses, form gyres in the central ocean basins north and
216 south of the equator (Mann and Lazier 2006). Those of the subtropics circulate in an
217 anticyclonic motion, while those in boreal subpolar regions rotate in a cyclonic motion: there
218 are no clear subpolar gyres in the Southern Hemisphere.

219 Subtropical gyres transport large quantities of heat away from the equator in western
220 boundary currents. These currents meander and cut off large gyrating bodies of water forming
221 cold and warm-core rings. Since the oceanic waters of anticyclonic gyres are permanently
222 stratified, the surface waters tend to have low average levels of nutrients (Mann and Lazier
223 2006). It was previously assumed that the subtropical gyres were areas of low productivity
224 but the currents along with the rings have unique properties which are used by different
225 organisms for breeding and migration to higher productivity sites in the temperate latitudes
226 (Mann and Lazier 2006). The productivity of subtropical gyres is greatly increased by the
227 presence of eddies and Ekman transport from peripheral gyres (Mann and Lazier 2006). Due
228 to the seasonal influence of monsoon winds over the Indian Ocean the circulation of the
229 Southern Indian Gyre is more variable than the Pacific and Atlantic subtropical Gyres
230 (Longhurst 1998).

231 Surface water circulation in the southern Indian Ocean is dominated by the Subtropical
232 Anticyclonic Gyre (Demopolous et al. 2003, Sultan et al. 2007). The gyre is largely located
233 on the western half of the Indian Ocean (Demopolous et al. 2003, Sultan et al. 2007), with the
234 eastern part being blocked by the South East Indian Ocean Ridge (Rogers et al. 2009). The
235 western boundary of the gyre is associated with a strong southward transport of water
236 associated with the Agulhas Current (Rogers et al. 2009). There is no gyre in the northern
237 Indian Ocean owing the position of the Asian continental landmass, but winds there blow

238 from the northeast from November to February (winter monsoon) and from the southeast
239 from May to September (summer monsoon) (Shankar et al. 2002). A transition period with
240 weak winds occurs between March to April and October (Shankar et al. 2002). A seasonally
241 reversing circulation in the upper north Indian Ocean is caused by the strong winds
242 experienced during the summer monsoon (Shankar et al. 2002).

243 In response to the lack of knowledge about the Southern Indian Ocean Gyre, a series of
244 surveys were undertaken from 2015 through to 2020 across the southern part of the Indian
245 Ocean in support of the IIOE-2. The first survey began in 2015 on-board the *Dr Fridtjof*
246 *Nansen*. The survey, in preparation for the 2nd International Indian Ocean Expedition (IIOE-
247 2), hosted scientists and technicians from six Western Indian Ocean countries. The first leg of
248 the survey, which collected data on the Southern Indian Ocean Gyre, began in Jakarta,
249 Indonesia, and ended in Port Louis, Mauritius (Groeneveld and Koranteng 2017). Preliminary
250 results from the survey regarding the anticyclonic circulation of the southern Indian Ocean
251 (Subtropical Indian Ocean Gyre) indicated that the gyre consisted of a number of smaller
252 eddies which impacted the flow regime and productivity within the gyre (Toresen et al.
253 2015). The preliminary results suggest that there was an increase in surface chlorophyll
254 within these eddies but the processes driving this are unknown (Toresen et al. 2015). The
255 anticyclonic circulation yielded low biological production within the gyre, but relatively
256 higher biological production away from the centre and towards the edges of the gyre
257 (Groeneveld and Koranteng 2017). Dufois et al. (2017) later hypothesized several processes
258 responsible for enhancing surface chlorophyll in the eddies in the South Indian Ocean.

259 The majority of the previous studies done on zooplankton communities across oceans have
260 been of a latitudinal rather than a longitudinal nature (e.g. McGowan and Walker 1979, 1985,
261 de Puelles et al. 2019) and have tended to focus on the change in community structure with
262 depth rather than with longitude. de Puelles et al. (2019) showed that the abundance of
263 zooplankton decreased with depth in the subtropical and tropical latitudes across the Atlantic,
264 Pacific and Indian Oceans. This decrease in abundance was noted in all oceans where the
265 influence of the surface layer is known to extend over 400 m with high zooplankton
266 abundance values found close to upwelling areas. de Puelles et al. (2019) concluded that
267 depth rather than oceans or biogeographical provinces was the primary factor structuring the
268 habitat and communities of zooplankton. McGowan and Walker (1979), in their study of the
269 structure of the copepod community of the North Pacific Central Gyre, showed that the only
270 group of species with a clear, general, seasonality of abundance was the most shallow-living

271 group and that members of this group live in a zone of strong temperature seasonality but
272 with no apparent seasonal change in primary production.

273 Previous studies done on the abundance and diversity of zooplankton in the Central South
274 Pacific Gyre (González et al. 2019) and the South Atlantic Gyre (Piontkovski et al. 2003)
275 have shown that the zooplankton biomass decreased towards the centre of both (Piontkovski
276 et al. 2003, González et al. 2019). The results from González et al. (2019) showed that there
277 was a significant segregation of the zooplankton communities divided into four zones
278 (eutrophic, mesotrophic, oligotrophic and the ultra-oligotrophic) with the gyre being the
279 ultra-oligotrophic zone.

280 Within the Southern Indian Ocean Gyre, mesopelagic fish densities were low but tended to
281 follow biological production. The highest density of plastics or micro-plastics was collected
282 along the gyre's edges and the majority of plastics were derived from the eastern edge.
283 Plastics were present in the majority of the water samples (Groeneveld and Koranteng 2017).
284 The presence of plastics within the Indian Ocean, in particular the ocean surface, sea-air
285 interface, threatens the existence of many marine organisms, in particular neustonic
286 zooplankton (zooneuston) living at/on the surface of the ocean.

287 Zooplankton are animals that “float” in the water column. They are the primary consumers of
288 the oceans and comprise many different animal groups such as crustaceans, molluscs,
289 cnidarians, and even fish in larval form. Zooplankton occur in a variety of sizes, from tiny,
290 microscopic ciliates to enormous jellyfish (Naumann 1917, Collard et al. 2015).

291 Holoplankton are zooplankton which remain on the surface of the ocean throughout their
292 lifecycle and comprise mostly of floating and rafting animals (e.g. Hydrozoa, Siphonophorae,
293 Gastropoda, Copepoda, Ostracoda, Euphausiacea, Chaetognatha) whereas meroplankton are
294 the zooplankton which comprise of individuals which have life stages at both the surface and
295 benthos of the ocean (e.g. Stomatopoda, Decapoda, fish larvae and fish eggs) (Santos et al.
296 2019). Zooneuston, hereafter simply referred to as neuston, are those animals that inhabit the
297 sea-air interface whereas zooplankton inhabit the subsurface layer (Naumann 1971, Collard et
298 al. 2015). The surface layer is a very important niche area for feeding and breeding of fish; in
299 particular commercially valuable fish species (Cardinale et al. 2003, Pusineri et al. 2005,
300 Collard et al. 2015).

301 As already noted, neuston comprises a variety of animals such as crustaceans (copepods,
302 mysids, euphausiids, amphipods, isopods, Ostracods and Decapod larvae), fish larvae,

303 polychaetes and their larvae, a wide variety of molluscs and their larvae, cnidarians
304 (hydromedusae, siphonophores and scyphozoans), insects (water striders), protostomes
305 (chaetognaths) and tunicates (salps, doliolids and appendicularians) (Marshall and Burchardt
306 2005). Due to their position at the sea-air interface, neuston are exposed to many
307 environmental and oceanographic stressors including high light intensity, wind stress,
308 turbulence and high (and low) temperatures, as well as high levels of predation from above
309 and below (Collard et al. 2015, Olivar et al. 2015). Most neustonic species have evolved
310 morphological and/or ecophysiological adaptations to this environment, which makes them
311 different from more typical zooplankton (Collard et al. 2015). Hydrozoans such as *Velella*
312 *velella* and *Physalia physalis* have a large part of their bodies occurring above the sea
313 surface, and they have adapted to their neustonic lifestyle by being able to withstand
314 prolonged desiccation and exposure to direct solar radiation (Liss and Duce 2005). Smaller
315 neustonic organisms are less susceptible to sinking and are able to float on the surface of
316 water due to the cohesive forces between water molecules which provide surface tension,
317 allowing them to float (Marshall and Burchardt 2005).

318 The stresses of life at the sea surface mean that comparatively few species occupy the
319 neuston permanently and choose rather to inhabit this ecotone temporarily (Olivar et al.
320 2015). Neuston can be categorised into three major ecological groups; euneuston (consisting
321 of epineuston and pleuston), facultative neuston and pseudoneuston (Hempel and Weikert
322 1972). The euneuston are adult organisms permanently inhabiting the neustonic realm, day
323 and night, and are most abundant in the direct vicinity of the ocean surface (Holdway and
324 Maddock 1983). Examples of euneuston are species of *Halobates* (water striders) and
325 floating and rafting animals (or pleuston), which have submerged bodies and aerial flotation
326 devices (e.g. *Velella velella*, *Porpita porpita* and *Physalia physalis*). Facultative neuston
327 inhabit the surface of the ocean temporarily, during part of their diel vertical migration
328 (Hempel and Weikert 1972, Holdway and Maddock 1983). Examples of facultative neuston
329 are neustonic larvae and juvenile stages of planktonic, nektonic or benthic species (Holdway
330 and Maddock 1983).

331 During diel vertical migration, zooplankton and fish move up and down the water column in
332 generally synchronised movement. Generally, plankton migrate towards the surface waters at
333 dusk and descend back to depth at dawn, with reverse patterns occurring as well (Brierley
334 2014). Neuston use diel vertical migration to escape predators that feed using sight and

335 possibly to avoid ultraviolet radiation due to its damaging effects. Metabolic clocks might
336 maintain diel vertical movement in the absence of light for example in the deep sea, cloud
337 cover, eclipse etc. (Brierley 2014).

338 Pseudoneuston occupy a depth range that extends to deeper layers: they are not confined to
339 the near-surface environment. Rather, the surface represents the uppermost extent of their
340 distribution (Holdway and Maddock 1983). The occurrence of neuston is dependent on the
341 time of day since a large number of animals enter the surface waters only during the night
342 (Hempel and Weikert 1972). Further classification of neuston can be made separating them
343 into epineuston and hyponeuston. Epineuston are those individuals living on the upper side of
344 the surface film of water or subaerially and are solely represented by the genus *Halobates*
345 (Cheng 1973, Holdway and Maddock 1983). Hyponeuston are those individuals living on the
346 lower side of the surface film of water extending to a depth of 5 cm (Marshall and Burchardt
347 2005, Gladyshev and Malyshevskiy 1982).

348 Euneuston and pleuston occurring in subtropical and tropical surface waters are well adapted
349 due to the unfavourable environmental conditions of irradiance, but environmental
350 conditions are uniform year-round with reduced wave action compared to the Boreal areas
351 (Hempel and Weikert 1972). In Boreal areas irradiance is lower and mixing by wave action in
352 more intensive, this results in the differences between the biotypes of neuston and plankton
353 being much reduced. The pseudoneuston (and very few euneuston e.g., the copepod
354 *Anomalocera patersoni*) only avoid the uppermost layer of the sea at noon when light-
355 intensities are extremely high (Hempel and Weikert 1972).

356 The sea surface is unfortunately home for more than just animals, and a large amount of
357 anthropogenic debris is known to occur and accumulate here (Thiel and Gutow 2004). Moore
358 et al. (2001) reported that in the North Pacific, the highest abundances of plastic microlitter
359 were recorded from the North Pacific Central Gyre having over 330 000 items of microlitter
360 per km² (Thiel and Gutow 2004). These plastics are frequently ingested by seabirds, pelagic
361 filter feeders and, when they settle on the seabed, by many benthic deposit feeders (Vlietstra
362 and Parga 2002, Thiel and Gutow 2004, Duis and Coors 2016). A tremendous amount of
363 anthropogenic debris enters the ocean regularly, approximately between 5 to 13 million
364 tonnes each year (Geyer et al. 2017). Microplastics are synthetic organic polymers smaller
365 than 5 mm in diameter, which have been recognised since the 1970s (Duis and Coors 2016).
366 There is an abundance of microplastics in the upper 20 cm of the sea surface in the gyres of

367 the North and South Atlantic, the North and South Pacific and the Indian Ocean (Duis and
368 Coors 2016). The circulation caused by gyral currents serves as a retention mechanism,
369 preventing plastics from escaping and moving to mainland coasts (Moore et al. 2001). A
370 surface current modelling study showed that many particles can be retained in a gyre for at
371 least 12 years (Inngraham and Ebbesmeyer 2000, Moore et al. 2001). The accumulation of
372 microplastics in the ocean, over time, ranges from low background levels to localized
373 “hotspots”, with gyres being particular hotspots for microplastics (Goldstein et al. 2012,
374 Wright et al. 2013). With an increase in the human population and consequently increasing
375 pollution and fishing activity, the assemblages of neuston could undergo drastic changes
376 (Collard et al. 2015). These changes could have a negative effect on the underlying
377 ecosystem. Therefore, an understanding of neustonic ecology is very important (Collard et al.
378 2015).

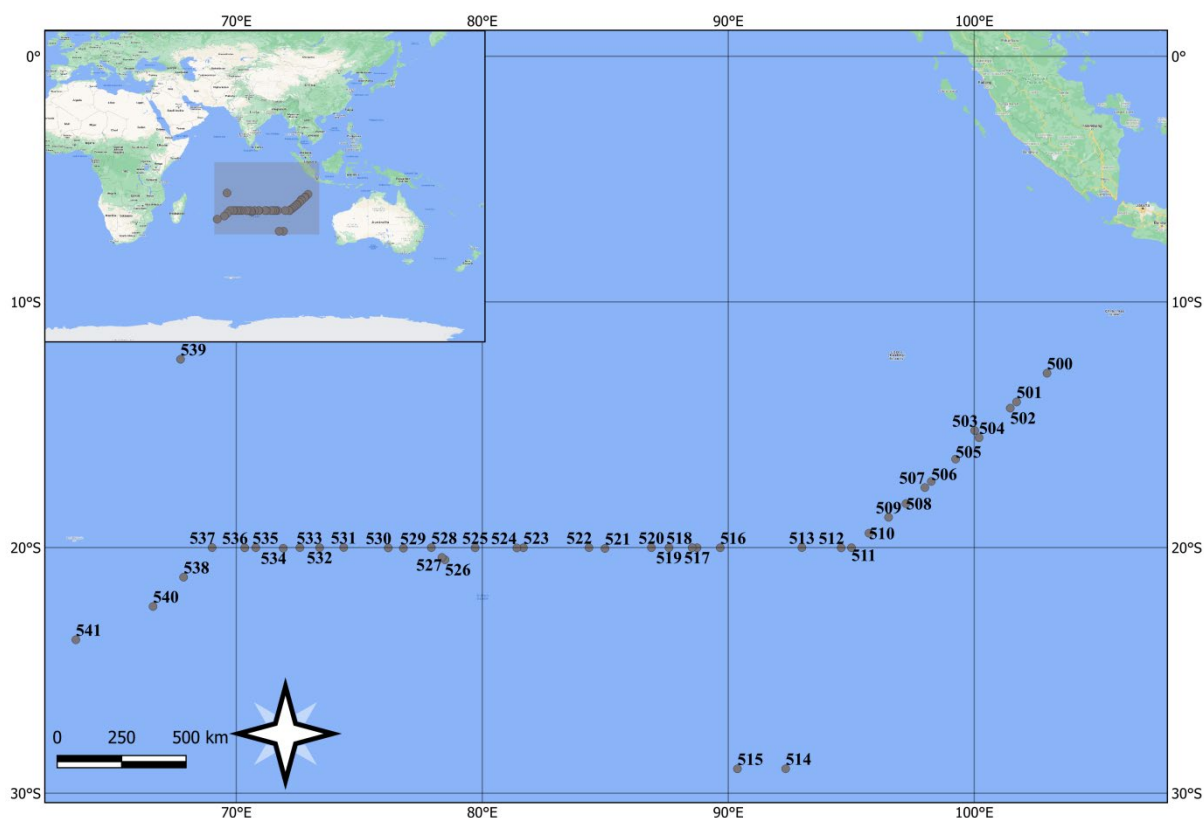
379
380 The aim of this study is two-fold. Firstly, it is to describe non-copepod neuston assemblages
381 across the Southern Indian Ocean Gyre and secondly it is to examine changes in assemblage
382 structure across the region, and to link such changes (where possible) to variations in the
383 oceanographic environment. It is hypothesised, following González et al. (2019) that
384 abundances will be lower in the centre of the gyre and that assemblages at the edge of the
385 gyre will be distinct from others. Following Hoeksema (2007) it is also anticipated that
386 assemblages closer to landmasses will comprise greater numbers of meroplanktonic taxa and
387 that there would be a higher abundance of neuston during the night than during the day
388 (Holdway and Maddock 1983).
389

390 **Materials and methods**

391

392 **Field sampling**

393 A total of 35 stations were sampled at regular intervals (100 nm, 150 nm) along a transect
394 line extending from Jakarta, Indonesia to Port Louis, Mauritius, between 26 June - 16 July
395 2015 (Figure 2) (Toresen et al. 2015). At each station, information on vertical changes in
396 temperature, salinity, oxygen and fluorescence were collected using a Seabird 911+ CTD
397 fitted with a Turner Design fluorometer.



398

399 **Figure 2:** Neuston stations sampled during the IOS leg 1 survey of the southern Indian Ocean
400 (QGIS version 3.16.3).

401

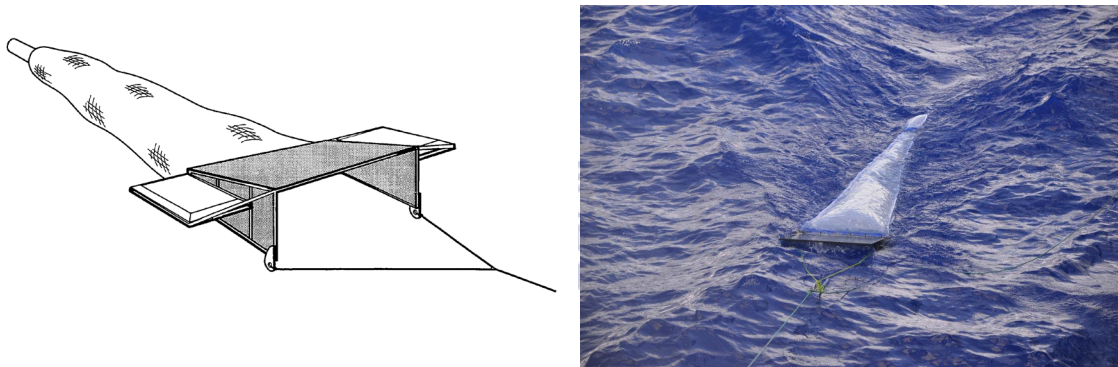
402 At each station, three neuston samples were collected using a rectangular manta trawl, with a
403 mouth frame measuring 19 cm high and 61 cm wide balanced by two wings and fitted with a
404 net having a 335 µm mesh and a 100 ml bottle fitted at the cod end (Figure 3). A mechanical
405 flow meter was fitted to the mouth of the net in order to determine the volume of water

406 filtered by the net. The manta net was only deployed in relatively calm weather and was
407 towed for a period of 15 minutes at a speed of 2-3 knots, horizontally at the ocean surface
408 (Toresen et al. 2015).

409 On retrieval, samples were immediately examined under a stereo microscope at various
410 magnifications, and all micro-plastics were removed. Thereafter the neuston was fixed and
411 preserved in either 96% ethanol or 4% buffered seawater formalin for later processing in the
412 laboratory (Toresen et al. 2015).

413

414



415

416

417 **Figure 3:** Diagram of a manta trawl, with a photograph showing how the manta trawl is used
418 to sample the sea-surface environment (Sandknop et al. 2002, Toresen et al. 2015).

419

420 Laboratory methods

421 Prior to microscopic examination, those samples in formalin were drained through a 180 μm
422 sieve and material was suspended in freshwater: samples in ethanol were examined without
423 rinsing. All samples were examined (and counted) without sub-sampling and specimens were
424 identified to morphospecies (Appendix A), except siphonophores, which were identified to
425 species level.

426 It is customary in ecological studies to identify sampled specimens to the lowest taxonomic
427 level using morphological characters. Ideally, this is to the level of species, as defined by the
428 Darwinian concept of a species (Darwin 1859, Mallet 2008, Pos et al. 2014). In practise,
429 however, this becomes difficult when samples are diverse and contain a large number of very

430 different taxonomic groups, because the expertise needed to identify all sample members to
431 species level is not available. In such circumstances, there is a tendency to focus on a
432 taxonomically distinctive and well-known taxon – a taxocene (Morin 2011). It is then
433 generally assumed that observations from the well-known can be extrapolated to the
434 unknown (Morin 2011). Although this assumption can sometimes be supported (e.g. Kuyper
435 et al., 2020), depending on the question being asked, there are instances when it cannot
436 (Morin 2011).

437 When attempts are made to identify everything to the lowest taxonomic level, individuals that
438 cannot be distinguished are often recorded as “indets”, and the number of indets will increase
439 as the dataset increases (Pos et al. 2014). Deciding whether to incorporate or ignore these
440 individuals in subsequence analyses can become tricky as they are associated with their own
441 errors. Consequently, there is little agreement among ecologists on how indets should be
442 treated and to what degree they may compromise the results (Pos et al. 2014).

443 An attempt is made here to be as inclusive as possible, and to use all the material in the
444 samples. However, given our lack of taxonomic expertise, it has been decided to use
445 morphospecies. As samples were processed, specimens with a readily identifiable
446 morphology were photographed, removed as vouchers and provided with unique labels, so
447 that it was possible to validate and cross-reference subsequent identifications. Indets are
448 consequently accorded taxon-level recognition because, with the exception of siphonophores,
449 all specimens are indets. It is acknowledged that in some cases, one morphospecies may
450 comprise more than one real species and it is understood that in other cases more than one
451 morphospecies could be assigned to a single species – especially in the case of larval
452 Decapods where different developmental stages may look quite different from one another.
453 However, the error is assumed to be consistent as only one person was involved in
454 morphospecies designation and sample analysis.

455

456 Data analyses

457 Temperature, salinity and chlorophyll/fluorescence were recorded in order to see if there
458 were any cohesive changes in the environment across the transect. Vertical profiles of the
459 temperature, salinity and chlorophyll/fluorescence were constructed using the computer
460 package Ocean Data View version 5.2.0. Even though neuston occur within the upper 20 cm
461 of the water column (Naumann 1971), the vertical profiles were constructed to a depth of 200

462 m in order to describe and assess the environment of migratory species. It should be
463 remembered that most of the samples were collected in an equatorial region, where
464 productivity is typically low at the surface and biomass generally peaks at some depth
465 (Stramma and Lutjeharms 1997).

466

467 In order to examine multivariate patterns in the assemblages sampled across the Southern
468 Indian Ocean Gyre, data were analysed using various software at the level of morphospecies,
469 (Appendix A). Patterns in community structure were assessed, *a priori*, by binning samples
470 into a) 5° longitudinal classes and b) Time of Day (day, dusk, night). The rationale behind
471 this was twofold. Firstly, our aim was primarily to determine if there were east-west changes
472 in the structure of neuston assemblages across the subtropical gyre. Owing to the patchy
473 nature of samples, it was necessary to bin them into 5-degree bins. And secondly, it is
474 understood that the composition and structure of neuston assemblages varies with time of day
475 (Hays 2003). Time of day was assessed from the sample log with reference to The NOAA
476 Solar Calculator (<https://www.esrl.noaa.gov/gmd/grad/solcalc/>) remembering that the ships
477 log was set as UTC (GMT): dawn and dusk were designated as the times of sunrise and
478 sunset, plus/minus 1 hr, respectively.

479 The biological data were first root-root transformed and a Bray-Curtis similarity matrix was
480 computed between samples. In order to determine if there was an effect of both longitude and
481 time of day on overall assemblage structure, data were analysed using a PERMANOVA
482 (permutational MANOVA), with the factors Longitude and Time of Day set as fixed and
483 random, respectively. Otherwise default settings were used: Type III sum of squares,
484 permutation of residuals under a reduced model and fixed effects sum to zero for mixed terms
485 (Anderson et al. 2008). These analyses were performed using PRIMER v7 + PERMANOVA
486 software (Clarke and Gorley 2015).

487 Given the significant effect of Time of Day on assemblage structure (see Results, below),
488 data were analysed separately for Day and Night samples; there being too few Dusk and
489 Dawn samples for analyses. Separate similarity matrices were constructed for Day and Night
490 samples, and the significance of a spatial pattern was tested using a one-way ANOSIM. A
491 matrix of pairwise R values between longitudinal bins was computed, which was visualised
492 using cluster analysis with group-average sorting. These analyses were conducted using
493 PRIMER v7 software (Clarke and Gorley 2015).

494 To test for differences in the average abundance of neuston across samples by Time of Day
495 (Day vs Night) and longitude, a two-way full factorial ANOVA was performed. Similar tests
496 were computed for the other univariate measures that summarise assemblages, namely
497 species richness and diversity (H'). Species diversity per sample was computed using the
498 Shannon Index (H'), following Krebs (2002). While sample averages were used to test for
499 changes in diversity across the transect as just outlined, all samples per longitudinal bin were
500 pooled to determine overall bin diversity, and error values around these were computed using
501 jack-knife methods (Magurran 2003).

502 The morphospecies responsible for 70% of the similarity (identity) of samples by Time of
503 Day and longitudinal bin were determined using a Similarity Percentage analysis (SIMPER).
504 Data were pooled by longitudinal bin and Time of Day, respectively, in these analyses owing
505 to the fact that there were too few data to generate robust patterns if they had been
506 partitioned. These analyses were conducted using PRIMER v7 software (Clarke and Gorley
507 2015).

508 A Distance Based Linear Model (DistLM) was computed to determine the environmental
509 variables driving community structure across longitudinal bins, by Day and Night separately.
510 The predictors used included seas surface temperature and salinity, the temperature and
511 salinity at 200 m, the depth of the upper mixed layer and of the fluorescence maximum as
512 well as integrated fluorescence and, the fluorescence value at the fluorescence maximum.
513 Marginal and sequential tests were performed by stepwise selection and the significance of
514 model outputs was assessed using adjusted R^2 . Owing to some fairly large gaps in the
515 environmental data (especially for fluorescence), analyses were confined **only** to those
516 samples with matching environmental data. The outputs of all DistLMs were visualised using
517 distance-based redundancy analysis (dbRDA) (Clarke and Gorley 2015).

518

519 **Results**

520

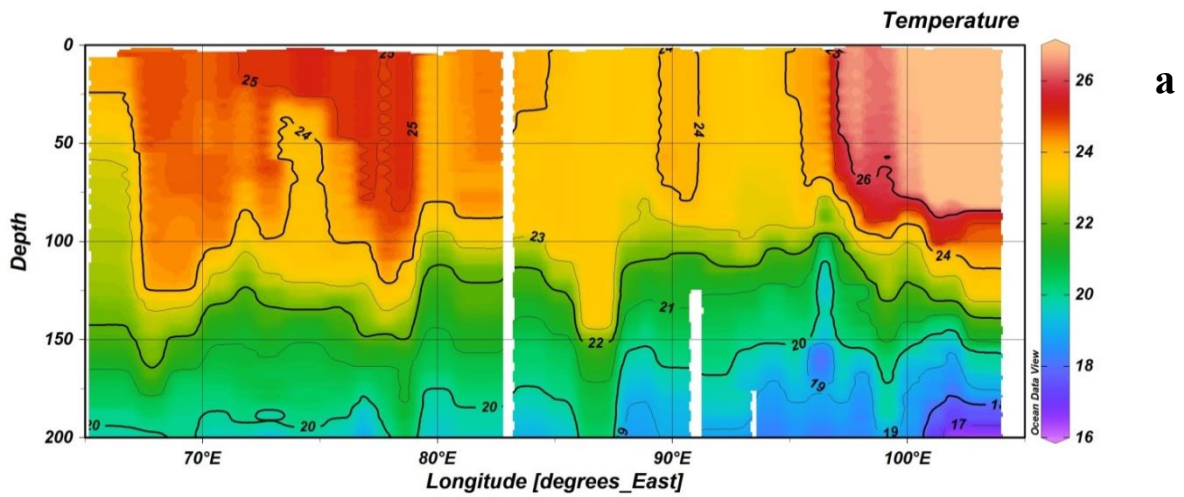
521 **Hydrography**

522 The vertical temperature profiles indicates that surface waters were generally warmer than
523 24°C across the transect, being warmest off Christmas Island to the east (Figure 4). Warmer
524 than surrounding waters were centred on the two anticyclonic gyres reported in the region at
525 ~ 76° E and ~88° E (Dufois et al. 2017). While the ocean temperature remained above 23°C
526 down to 100 m, an upward doming of cooler isotherms is noted east of ~90° E, reflecting the
527 cyclonic gyre noted in that region (Dufois et al. 2017). The lowest temperature at 200 m
528 (17°C) was recorded off Christmas Island.

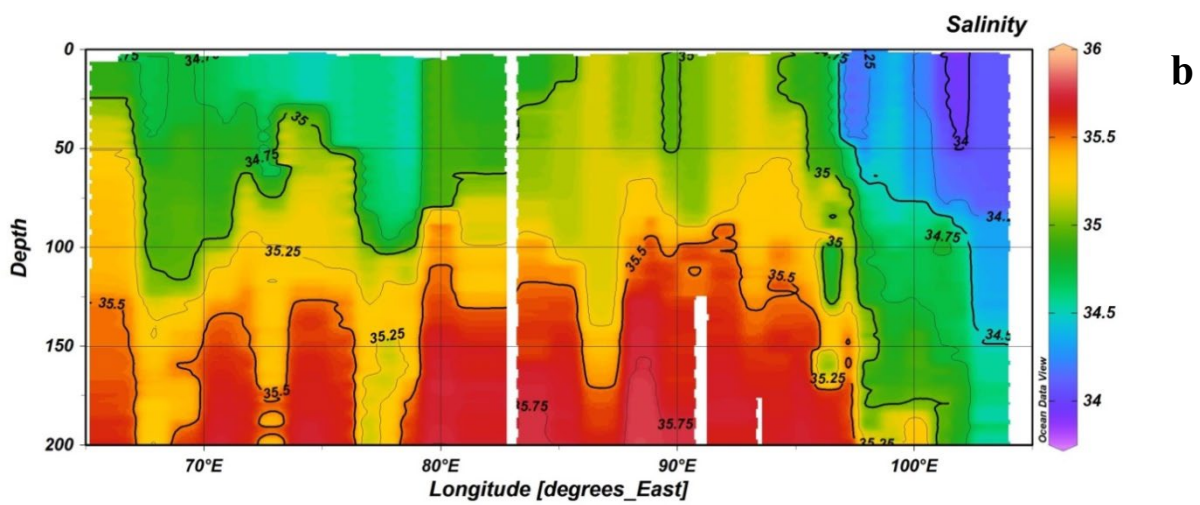
529 Salinity was lower at the surface than it was at depth across the width of the transect, being
530 lowest at the surface in the easternmost stations (Figure 4). Interestingly, the distribution of
531 surface salinities mirrored the distribution of temperatures, being slightly higher in cooler
532 than warmer waters. Indeed, there was a significant negative relationship between the two
533 variables (Pearson's $R=-0.94$, $p<0.05$). An apparent upward doming of high salinity water (>
534 35.5 psu) was evident to 100 m, west of ~100° E.

535 There was a relatively deep fluorescence maximum across the transect, centred at ~ 120 m,
536 except in the extreme west where it occurred at ~45 m (Figure 4). Dufois et al. (2017) noted
537 that the vertical position of the chlorophyll maximum closely tracked the mixed-layer depth,
538 hence it approximates the distribution of deeper isotherms and isohalines.

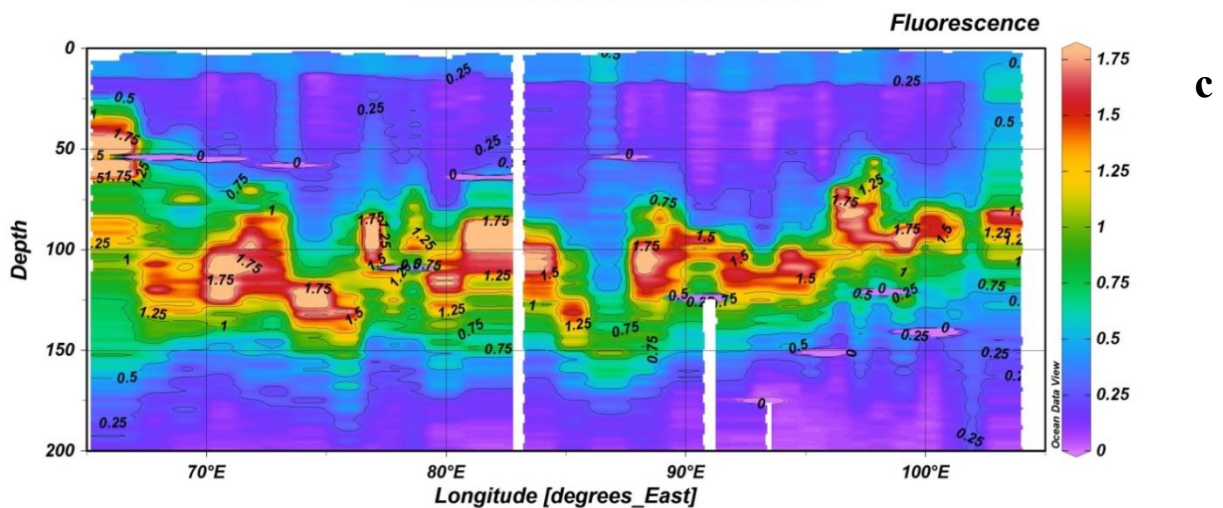
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543 **Figure 4:** Vertical profiles in temperature (a), salinity (b) and fluorescence (c) during the IOS
 544 leg 1 survey of the southern Indian Ocean (using Ocean Data View version 5.2.0(ODV)).

545

546 Neuston

547 Assemblage composition

548 Overall, the non-copepod neuston was dominated by Ostracods though mysids, juvenile
549 chaetognaths, siphonophores and amphipods were also common (Table 1). The greatest
550 number of morphospecies was observed amongst Amphipoda (19 Hyperiida, six Gammarida)
551 and Siphonophora (21 Calycophorae, three Physonectae, one Cystonectae). Gastropods were
552 represented by 20 morphospecies (17 Pteropoda, one Littorinimorpha, one Nudibranchia, one
553 Caenogastropoda); there were ten Thaliacea (eight Salpida, two Pyrosomatida, one Doliolida)
554 and seven each of Stomatopoda, fish larvae and Decapoda. The taxon that was most
555 frequently recorded across all samples was Siphonophora, which occurred in 83 of the 87
556 samples, followed by Gastropoda (72 samples), non-siphonophoran Hydrozoa (71) and
557 Ostracoda (70) (Table 1). Of the siphonophores, *Abylopsis eschscholtzii* was most abundant
558 followed by *Chelophyes contorta* and *Bassia bassensis*.

559 The majority of the specimens collected (88%), and the greatest variety of morphospecies,
560 were facultative neuston (Appendix A). Of the balance, ~10% comprised euneuston
561 (pleuston) *Porpita* and *Velella* (Hydrozoa), *Physalia* (Siphonophora) and epineuston, Water
562 Strider, *Halobates* (Insecta) and ~2% were hyponeuston [morphospecies of *Spirula*
563 (Cephalopoda), *Janthina* and *Glaucus* (Gastropoda)]. Although the relative proportions of the
564 different types of neuston varied little across the transect, facultative neuston were slightly
565 more common by night (~90%) than day (86%).

566 Holoplanktonic morphospecies dominated neuston by day (~84.5%) and night (85.5%). The
567 contribution by meroplanktonic morphospecies was greater at the longitudinal extremes of
568 the transect than it was at the centre of transect (Figure 5), especially by night.

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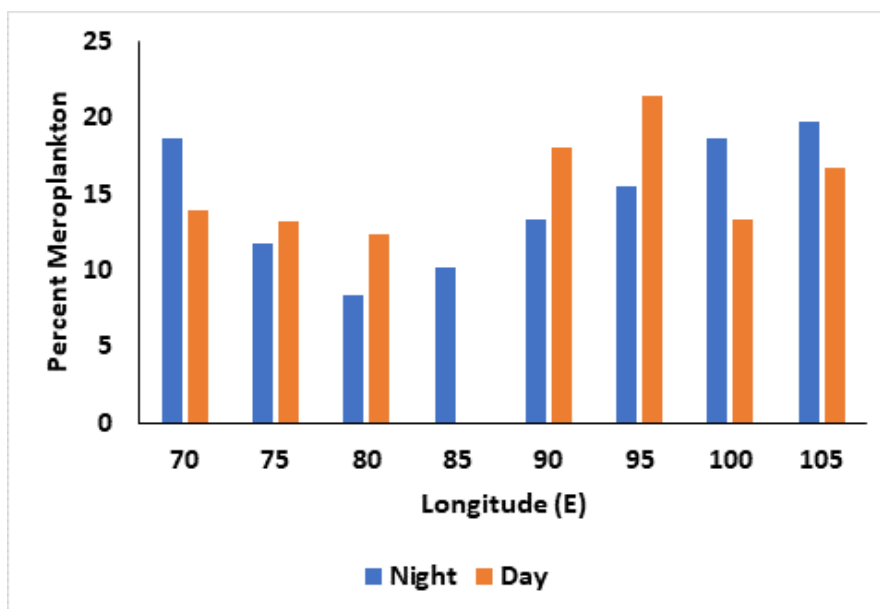
571 **Table 1:** The average abundance, diversity (number of morphospecies) and frequency of
 572 occurrence of the major taxa recovered in neuston samples across the Southern Indian Ocean
 573 Gyre of the Indian Ocean during 2015. Their relative rankings also shown.

Taxon	Average abundance	Rank	no morphospecies	Rank	no samples	Rank
Amphipod	27,28	5	25	1	70	5
Barnacle	0,06	22	3	14	4	22
Cephalopod	0,07	20	4	11	6	19
Chaetognath	50,15	3	1	18	65	7
Decapod	18,63	6	7	5	66	6
Doliolid	0,84	15	1	18	26	14
Euphausiid	1,78	13	2	16	5	20
fish eggs	1,00	14	1	18	28	13
Fish Juvenile	0,11	19	5	8	7	18
Fish larvae	1,84	12	7	5	50	10
Gastropod	10,47	7	20	3	72	2
Hydrozoa	9,70	8	4	11	71	3
Isopod	0,30	17	4	11	11	16
Mysid	59,11	2	1	18	57	9
Ostracod	80,03	1	5	8	71	3
Polychaete	0,33	16	5	8	11	16
Salp	2,40	10	10	4	60	8
Scyphozoa	0,07	20	2	16	5	20
Siphonophore	31,45	4	25	1	83	1
Sipunculid	0,01	23	1	18	1	23
Stomatopod	2,24	11	7	5	34	12
Water strider	4,46	9	3	14	47	11
Zooanthidae	0,28	18	1	18	19	15

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583 **Figure 5:** Changes in the relative contribution by meroplanktonic morphospecies to neuston
584 abundance across the Southern Indian Ocean Gyre in the Indian Ocean during 2015 with
585 longitude and time of day.

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Effects of longitude and time of day

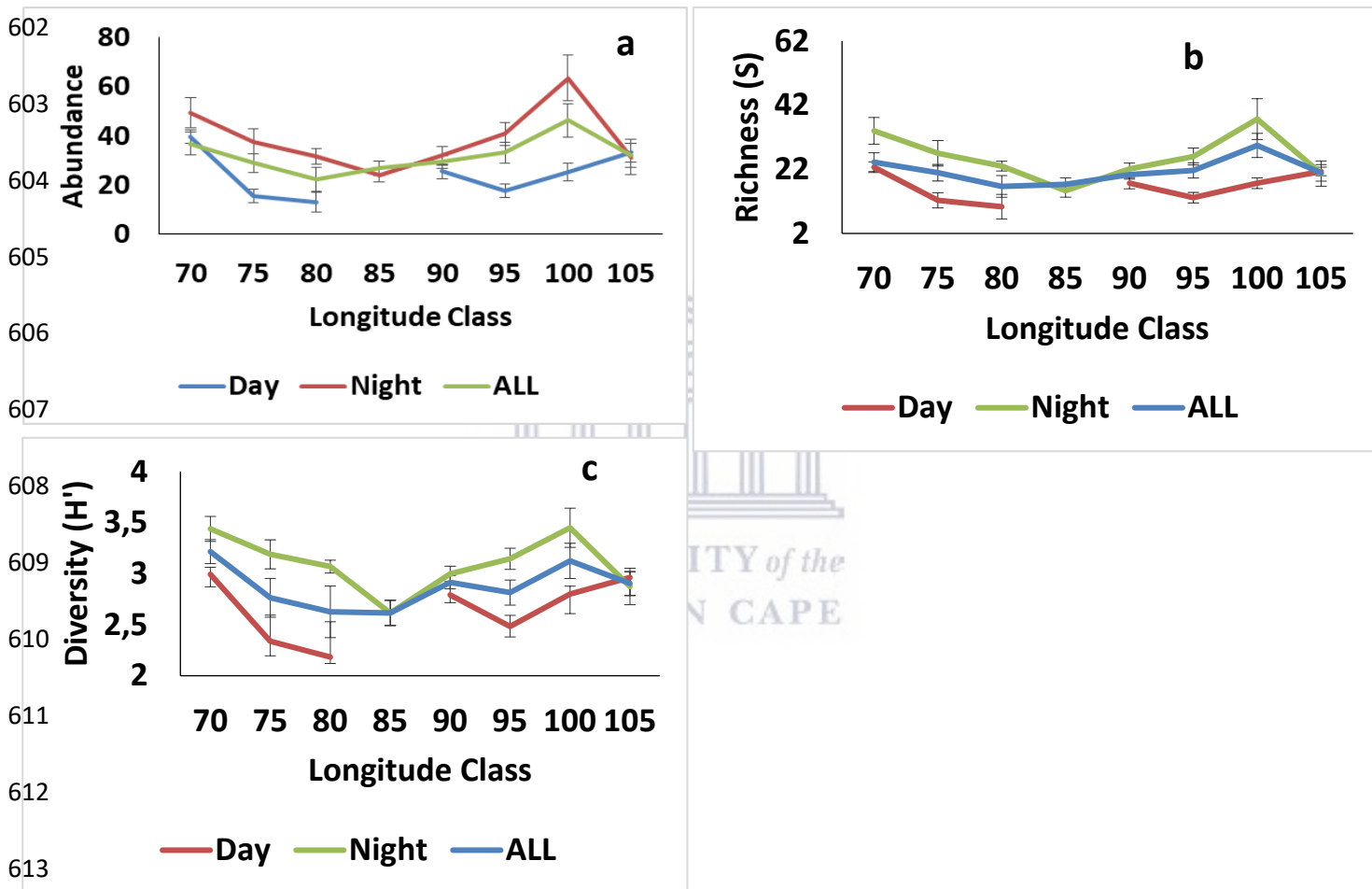
The results from the PERMANOVA routine indicated that both *a priori* factors influenced the structure of neuston assemblages across the transect, individually and interactively, despite the high residual sums of squares (Table 2). As a consequence, data have been treated separately by Time of Day in all subsequent analyses.

593 **Table 2:** PERMANOVA results to examine the effects of Time of Day and Longitude on the
594 structure of neuston assemblages across the IOS Leg 1 survey in the southern Indian Ocean.

Source	DF	SS	MS	Pseudo-F	P(perm)	No Unique Permutations	P(MC)	Estimate of components of variation
Longitude	7	30626	4375,1	1,495	0,042	998	0,029	157,67
Time of Day	3	17953	5984,4	4,653	0,001	998	0,001	316,89
Longitude x Time of Day	9	26668	2963,1	2,304	0,001	997	0,001	408,92
Residual	64	82308	1286,1					1286,1

595 The results of the two-way, full factorial ANOVA indicate that the total abundance of the
 596 neuston in the samples was significantly impacted by Time of Day ($F_{1,52}=34.49, p<0.0001$),
 597 Longitude ($F_{6,52}=4.38, p<0.005$) and the interaction between both factors ($F_{6,52}=3.01$,
 598 $p<0.05$). Abundance was higher during the night than the day, and tended to be higher at the
 599 extremes of the transect and lowest around 80° E. (Figure 6a)

600
601



614

615 **Figure 6:** Diel variability in the mean abundance (a), richness (morphospecies, b) and
 616 diversity (H' , c) of the non-copepod neuston collected across the Southern Indian Ocean Gyre
 617 in 2015. Error bars are standard error.

618

619 Effectively similar temporal and spatial results were observed in terms of sample diversity,
 620 either measured as richness or diversity (H' : Table 3, Figure 6b, c). Diversity was higher in
 621 samples collected at night than during the day, and was greater at the extremes of the transect
 622 than around 80° E.

623

624 **Table 3:** Effect of Longitude, Time of Day and Longitude and Time of Day on sample
 625 diversity and species richness.

Measure	Abundance		Species Richness		H'	
	F	p	F	P	F	P
Intersect	531,52	<0.0001	632,56	<0.0001	4132,74	<0.0001
Longitude	4,38	<0,005	2,97	<0,05	2,37	<0,05
Time of Day	34,49	<0,0001	38,38	<0,0001	32,52	<0,0001
Longitude and Time of Day	3,01	<0,05	2,30	<0,05	2,46	<0,05

626

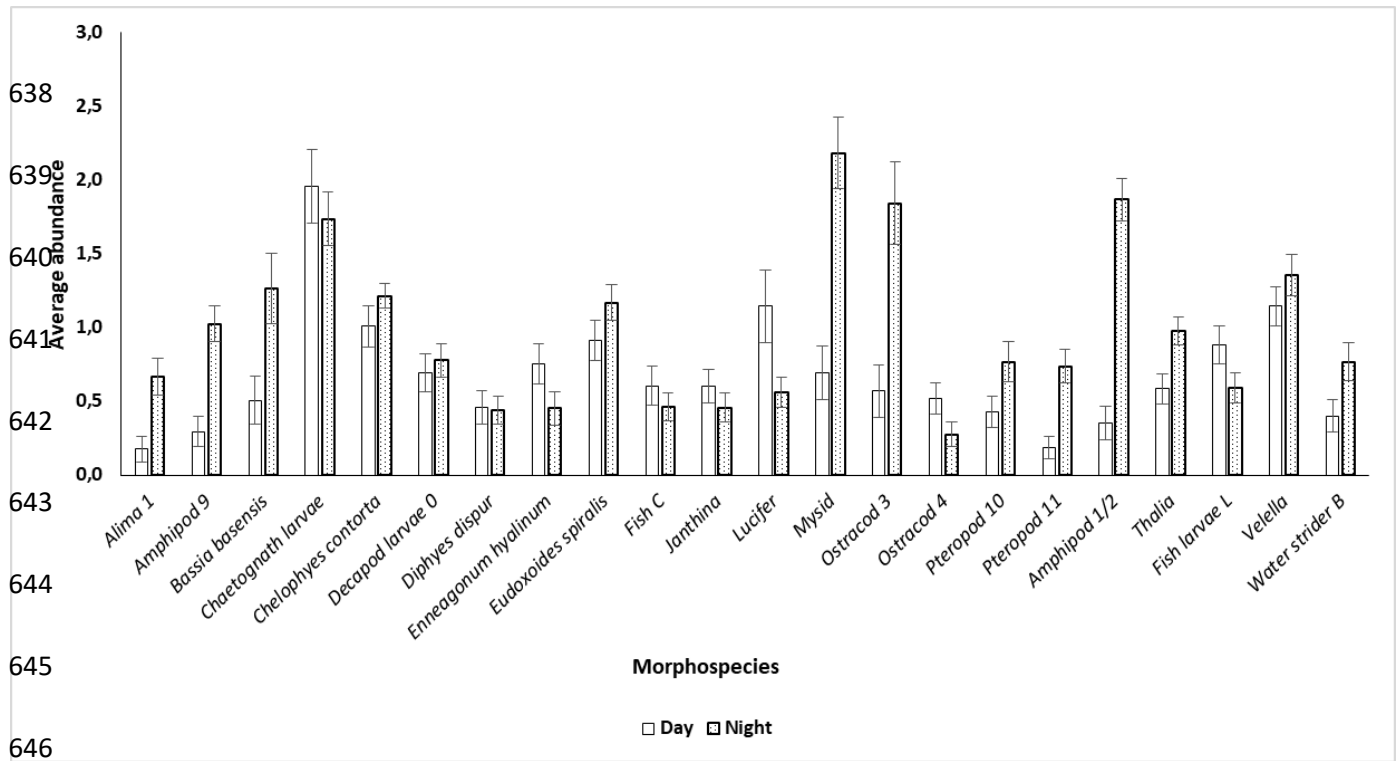
627 Composition

628 The average abundance (\pm se) of those morphospecies that were responsible for 70% of the
 629 dissimilarity (SIMPER analysis) between Day and Night samples (across all longitudinal bins
 630 are shown in Figure 7: those responsible for differences between longitudinal bins (across
 631 Time of Day) are shown in Figure 8.

632 As noted previously, the majority of taxa increased in abundance at night. With the exception
 633 of water striders, the other members of the euneuston (*Janthina*, *Velella*) did not differ much
 634 in abundance with Time of Day: only *Enneagonum hyalinum*, *Lucifer* and Ostracod 4 were
 635 more common during the day than at night.

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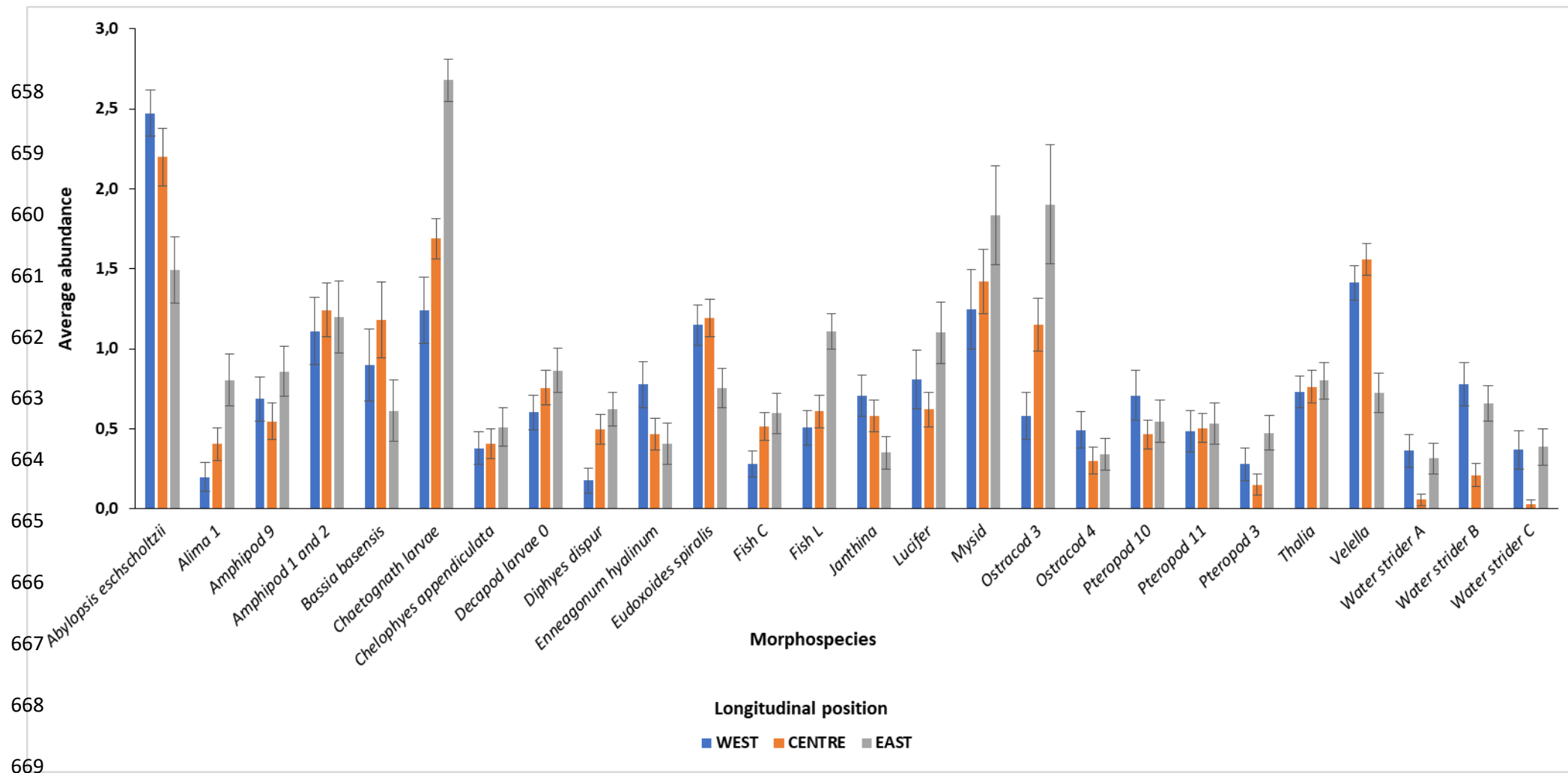
637



647 **Figure 7:** The average abundance of the morphospecies responsible for 70% dissimilarity
 648 between day and night samples collected in the IOS Leg 1 survey of the southern Indian
 649 Ocean. Data pooled by longitudinal bin. Error bars are standard error.

650

651 The data shown in Figure 8 indicate that all the taxa that were responsible for the identity of
 652 each longitudinal bin were found across the transect, though some did display a pattern of
 653 spatial distribution. For example, the siphonophores *Abylopsis eschscholtzii* and *Enneagonum*
 654 *hyalinum* were more common in the west (grids 70, 75, 80), whilst *Alima 1*, chaetognath
 655 larvae, mysids and Ostracod 3 were more abundant in the east (grids 95, 100, 105). Few taxa
 656 were most abundant in the central part of the transect (grids 80, 85, 90, 95), though water
 657 striders were notably uncommon there, by comparison with the extremes of the transect.



670 **Figure 8:** The average abundance of the morphospecies responsible for 70% dissimilarity between neuston samples collected in the centre (75 -
671 85° E), East (95 - 105° E) and balance of the IOS Leg 1 survey in the southern Indian Ocean. Data pooled by time of day bin. Error bars are
672 standard error. .

673 Assemblage structure

674 The results of the ANOSIM indicate that there was a significant difference in the structure of
675 assemblages across the transect, by Day (global $R=0.381$, $p<0.005$), Night (global $R=0.413$,
676 $p<0.005$) and overall (global $R=0.229$, $p<0.005$). The cluster analyses generated from the
677 matrices of pairwise R values are shown in Figure 9 and reveal some interesting patterns. The
678 night-time samples indicate that assemblages between 80 - 85° E are very distinct from the
679 balance, and that those at the extremes of the transect (between 65 - 75° E; 95 - 105° E) are
680 more similar to each other than to those immediately surrounding 85° E (Figure 9a). In the
681 case of the samples collected during the day, those between 75 - 80° E were very different
682 from the balance: no samples were collected during the between 80 - 85° E. The rest of the
683 samples fall into two groups that span 75 - 80° E, with those collected between 100 - 105° E
684 being dissimilar (Figure 9b). Examination of the cluster analysis for all data indicate that
685 samples from between 75 - 85° E were more similar to each other than to the rest of the
686 samples, which otherwise clustered out into groups that were approximate longitudinal
687 opposites of each other (Figure 9c).

688 The results of the cluster analyses shown in Figure 9 are graphically re-interpreted in Figure
689 10. Although there are some differences in the way that the different longitudinal bins relate
690 to each other by Time of Day, in general it would appear that there is a symmetry to the
691 pattern either side of a “core” between 75 - 85° E.

692 The results of a SIMPER analysis exploring the taxa that are responsible for similarities
693 within, and differences between the three clusters identified in Figure 9c are graphically
694 illustrated in (Table 6). From this it can be seen that assemblages between 75 - 85° E are
695 characterised by siphonophores and few mesozooplankton. Amphipods, chaetognaths,
696 mysids, water-striders and *Lucifer* were more common at the other stations, with those to the
697 east (95-105° E) supporting greatest numbers of meroplanktonic crustaceans.

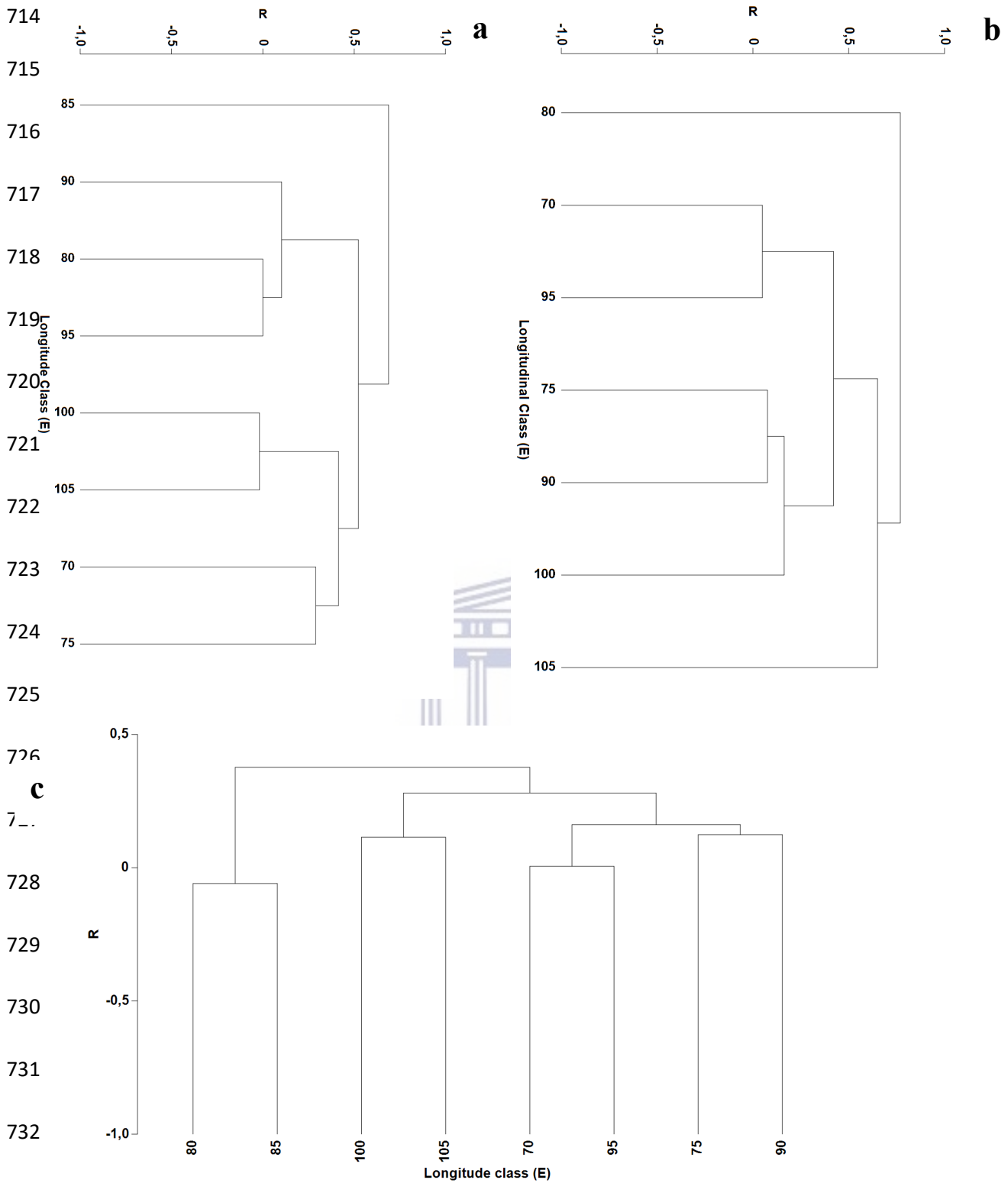
698

699 Owing to their high richness and widespread distribution across the region, the
700 siphonophores, amphipod and gastropod data can be analysed separately. In the case of the
701 former, there was no overall diel (ANOSIM $R=0.07$, $p>0.05$) or longitudinal (ANOSIM $R=-$
702 0.09 , $p>0.05$) change in assemblage structure. Although there was also no distinct
703 longitudinal pattern in amphipod assemblage structure across the transect (ANOSIM $R=0.09$,
704 $p>0.05$), assemblages were different, by day and night (ANOSIM $R=0.29$, $p<0.005$), being

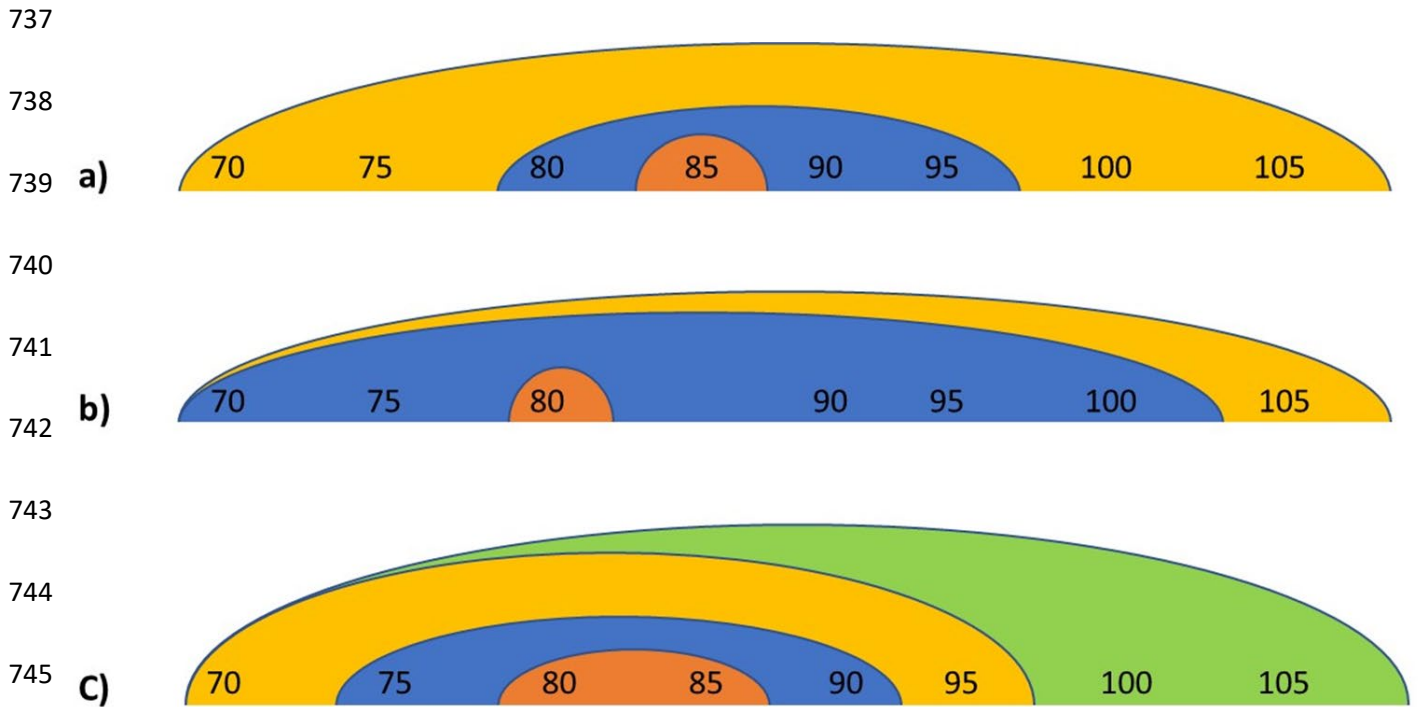
705 both more abundant (13 vs 4 per haul, respectively) and diverse (20 vs 12 morphospecies) at
706 night. Although there was no apparent longitudinal pattern to the distribution of assemblages
707 of neustonic gastropods (ANOSIM $R=0.11$, $p>0.05$), there were significant diel differences in
708 assemblage structure (ANOSIM $R=0.34$, $p<0.005$). Overall, day time samples were less
709 species rich than night-time samples (12 vs 18 morphospecies, respectively), and contained
710 fewer individuals (4.6 vs 33.4, respectively). Both morphospecies of *Creseis* were more
711 common by day than night, whilst heteropods (*Atlanta* and Pterotracheidae) were more
712 common at night.

713





735 **Figure 9:** Cluster analyses showing the significant difference in the structure of assemblages
 736 across the transect during the night-time (a), day-time (b) and all data (c).



748
 749 **Figure 10:** Graphical re-interpretation of cluster analyses shown in Figure 9 for night-time
 750 (a), day-time (b) and all samples collected during IOS Leg 1 survey of the southern Indian
 751 Ocean.

752
 753 The results of the DistLM are shown in Tables 4 and 5, for Day and Night samples,
 754 respectively. The results of the marginal tests indicate that none of the predictors associated
 755 with fluorescence were significantly related to the longitudinal structure of the day-time
 756 assemblages (Table 4a). That said, three (depth of the fluorescence maximum, sea surface
 757 fluorescence and integrated fluorescence) were included in the final model ($R^2=0.49$, adjusted
 758 $R^2=0.33$; Table 4b). The dbRDA plot for the day-time samples is shown in Figure 11, the first
 759 two axes of which explain 56% of fitted and 27% of total variation.

761 **Table 4a:** Marginal test results of the DistLM exploring the effect of different environmental
 762 variables on the structure of day-time neuston assemblages observed across the IOS Leg 1
 763 survey in the southern Indian Ocean.

Variable	SS(trace)	Pseudo-<i>F</i>	<i>P</i>	Prop.
Latitude	4001,5	2,180	0,006	0,072
Longitude	4075,8	2,223	0,01	0,074
Sounding	1157,3	0,597	0,876	0,021
SST (sea surface temperature)	3551,3	1,918	0,033	0,064
Temperature at 200 m	4952,4	2,748	0,001	0,089
SSS (sea surface salinity)	3490,7	1,883	0,029	0,063
S (salinity) at 200 m	4242,1	2,321	0,007	0,077
SSFI (Sea surface fluorescence)	3189,9	1,710	0,055	0,058
D FI max (deep fluorescence maximum)	2247	1,184	0,259	0,041
FI at Chl max (fluorescence at chlorophyll maximum)	3126,5	1,674	0,068	0,056
Integr FI (Integrated fluorescence)	3244,1	1,741	0,061	0,059

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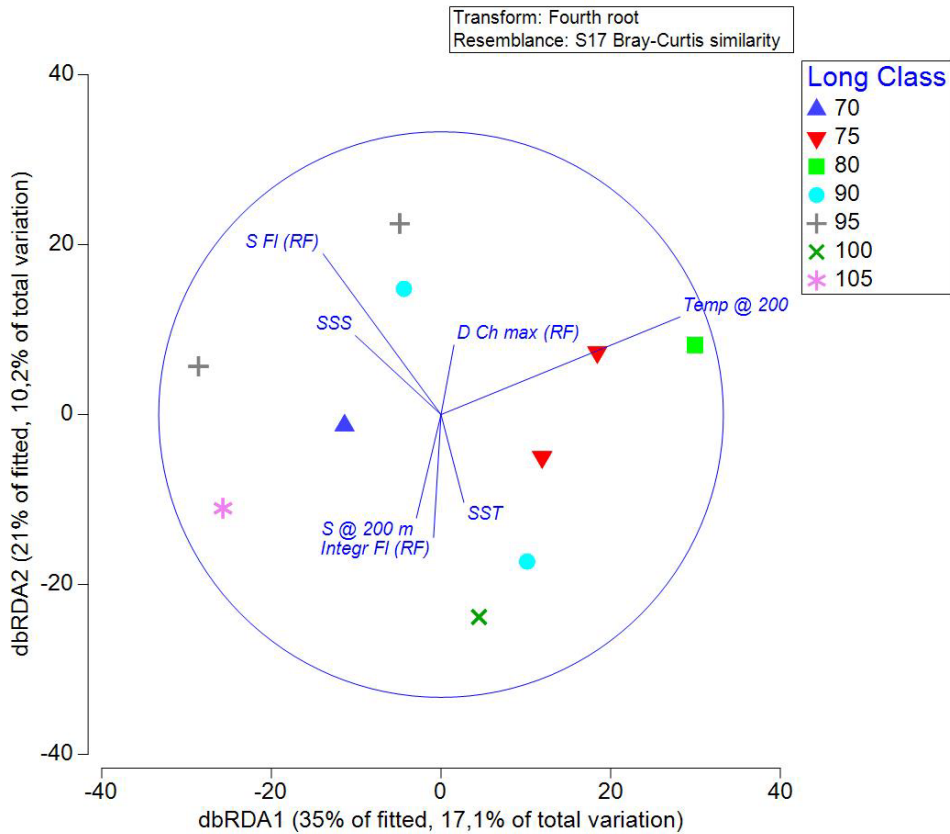
765 **Table 4b:** Sequential test results of the DistLM exploring the effect of different
 766 environmental variables on the structure of day-time neuston assemblages observed across
 767 the IOS Leg 1 survey in the southern Indian Ocean. See Table 4a for full labels to variables.

Variable	Adj R²	SS(trace)	Pseudo-<i>F</i>	<i>P</i>	Prop.	Cumul.	res.df
+Temp at 200 m	0,057	4952,4	2,7483	0,002	0,089	0,089	28
+S at 200 m	0,111	4607,5	2,7133	0,002	0,083	0,173	27
+D FI max	0,155	3882,3	2,4053	0,005	0,070	0,243	26
+SSS	0,206	4019,3	2,648	0,003	0,073	0,315	25
+Integr FI	0,237	2946,9	2,0207	0,023	0,053	0,368	24
+SSFI	0,266	2739,1	1,9528	0,033	0,049	0,418	23
+SST	0,324	3837,9	2,9706	0,002	0,069	0,487	22

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772 **Figure 11:** dbRDA plot illustrating the relationship between environmental predictors and
773 the structure of day-time neuston assemblages observed across the IOS Leg 1 survey in the
774 southern Indian Ocean.

775

776 Interestingly, and perhaps tellingly, in the case of the night- time samples, only the depth of
777 the fluorescence maximum was not significantly related to the structure of assemblages along
778 the transect; both measures of chlorophyll biomass were significant (Table 5a). That said,
779 only the depth of the fluorescence maximum was included in the full, sequential model
780 ($R^2=0.62$, adjusted $R^2=0.45$; Table 5b). The first two axes of the dbRDA plot (Figure 12)
781 explain 61% of fitted and 38% of total variation. Note the lesser number of stations included
782 in the night- time than day time model owing to gaps in the environmental data (Fig 11 vs
783 Figure 12).

784

785 **Table 5a:** Marginal test results of the DistLM exploring the effect of different environmental
 786 variables on the structure of night-time neuston assemblages observed across the IOS Leg 1
 787 survey in the southern Indian Ocean. See Table 4a for full labels to variables.

788

Variable	SS(trace)	Pseudo-<i>F</i>	<i>P</i>	Prop.
Latitude	3586,5	2,233	0,009	0,092
Longitude	6640,5	4,525	0,001	0,171
Sounding	4785	3,083	0,001	0,123
SST	6258,7	4,215	0,001	0,161
Temp at 200 m	3986	2,510	0,006	0,102
SSS	5779,5	3,836	0,001	0,148
S at 200 m	5483,9	3,607	0,001	0,141
SSFI	3472,5	2,155	0,016	0,089
D FI max	1995,3	1,189	0,254	0,051
FI at Chl max	5188,9	3,383	0,001	0,133
Integr FI	4315,3	2,743	0,002	0,111

789

790 **Table 5b:** Sequential test results of the DistLM exploring the effect of different
 791 environmental variables on the structure of night-time neuston assemblages observed across
 792 the IOS Leg 1 survey in the southern Indian Ocean. See Table 4a for full labels to variables.

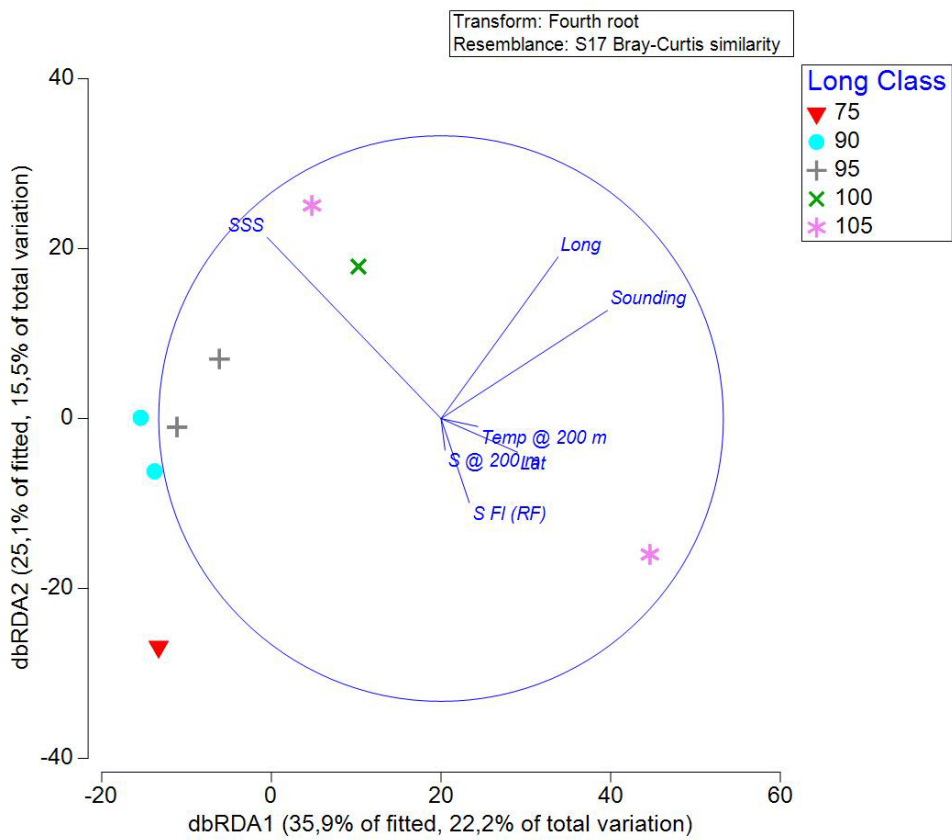
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Variable	Adj R²	SS(trace)	Pseudo-<i>F</i>	<i>P</i>	Prop.	Cumul.	res.df
+Longitude	0,133	6640,5	4,525	0,001	0,171	0,171	22
+SSS	0,205	4023,3	2,990	0,002	0,103	0,274	21
+ D FI max	0,248	2798,5	2,198	0,008	0,072	0,346	20
+SSFI	0,321	3628,3	3,157	0,001	0,093	0,439	19
+Latitude	0,382	3020,3	2,889	0,001	0,078	0,517	18
+Sounding	0,424	2230,3	2,286	0,014	0,058	0,574	17
+Temp at 200 m	0,452	1757,6	1,896	0,04	0,045	0,619	16

794

795

796



797

798 **Figure 12:** dbRDA plot illustrating the relationship between environmental predictors and
799 the structure of night-time neuston assemblages observed across the IOS Leg 1 survey in the
800 southern Indian Ocean.

801

802 Discussion

803

804 Vertical profile for environmental data

805 During the austral winter, the Subtropical Gyre in the southern Indian Ocean is generally
806 characterised by anticyclonic eddies and elevated surface chlorophyll concentrations
807 (references in Dufois et al. 2017). In their study, which was conducted on the same cruise as
808 the present data were collected, Dufois et al. (2017) used Bio-Argo floats and *in situ*
809 sampling to understand the mechanisms responsible for this chlorophyll enrichment. They
810 noted the presence of three eddies: one cyclonic eddy near 89° E and two anticyclonic eddies
811 at approximately 78.5° E and 87° E (Dufois et al. 2017). The cyclonic eddy was the smallest
812 of the three and had a lifespan of 70 days, while the anticyclonic eddies had a lifespan
813 between 117 - 132 days (Dufois et al. 2017).

814 Dufois et al. (2017) concluded that the anticyclonic eddies “can trap productive waters
815 through horizontal advection and/or locally modify the chlorophyll distribution through
816 enhanced convective mixing” (p 3262). They noted that the warm waters in the west,
817 associated with the largest of the eddies, were drawing in waters from the north, which
818 themselves originated in part from the Indonesian Throughflow and had been propagated
819 westward – hence the lower salinity. Usually, higher sea surface temperatures are associated
820 with high levels of evaporation and salinity but south of the equator, Indonesian Throughflow
821 Water transports freshwater from Indonesian seas to the south eastern tropical Indian Ocean,
822 which maintains a low Sea Surface Salinity (SSS) (Du and Zhang 2015). The anticyclonic
823 movement of the Indian Ocean Gyre recirculates the low salinity water and maintains a
824 negative SSS south of the equator (Du and Zhang 2015).

825

826 Neuston

827 On balance, the neuston communities examined here are broadly similar to those of previous
828 studies, as recently reviewed by Helm (2021). That said, most of the work that has been
829 conducted on neuston has tended to focus on specific taxonomic groups, often with a focus
830 on fish larvae (e.g. Davenport and Rees 1993), and comprehensive research on regional
831 communities as a whole are scant. Zaitsev (1970) reviewed the work that he and colleagues
832 had led in the Black Sea and other areas; Hempel and Weikert (1972) studied neuston in the
833 North Atlantic Ocean, whilst Holdway and Maddock (1983) reported on the neuston
834 collected along a transect extending from Indonesia, across the Indian Ocean and into the

835 Mediterranean Sea. Brodeur et al. (1987) briefly examined the neritic neuston of the NE
836 Pacific, with reference to the salmon resources there, while Echelman and Fishelson (1990)
837 and Mantha et al. (2019) have documented communities in the Red Sea. Otherwise
838 comprehensive studies that can be used for comparative purposes have been hard to find. For
839 example, a recent study of 54 neuston samples collected in the NE Pacific, taken to
840 contextualise the fauna of the Great Pacific Garbage Patch, failed to report the presence of
841 many key neuston groups (Egger et al. 2021) and cannot be used here.

842

843 Common taxa.

844 The eueneuston collected here is broadly similar to that noted by others and comprises
845 floating and rafting animals that permanently inhabit the surface film of the ocean (Weikert
846 1970, Grave 1971, Zaitsev 1971, Hempel and Weikert 1972). These animals do not display
847 diel vertical migration and are constantly on the ocean surface regardless of time of day
848 (Weikert 1970, Grave 1971, Zaitsev 1971, Hempel and Weikert 1972). That said,
849 morphospecies of *Halobates* were more common by night than by day. The two species of
850 *Halobates* that inhabit the Indian Ocean are *Halobates micans* and *Halobates germanus* both
851 of which occur in water with a temperature greater than 25°C (Harada 2005). *Halobates*
852 *micans* is the only species that can be found in all three ocean basins, while *H. germanus* is
853 widely regarded as occurring mostly in coastal waters (Cheng 1985; Harada et al. 2013). It is
854 suggested here that the lower abundance of *Halobates* in day-time trawls reflects net
855 avoidance and night blindness, as these animals detect food by sight or via the sensory
856 detection of water ripples created by struggling prey at the ocean surface (Sagaydachnyy
857 1975; Cheng 1985), suggesting a sensory disadvantage during the night.

858 As noted earlier, the other species/morphospecies that were slightly more common by day
859 than night were *Enneagonum hylinum*, *Lucifer* and Ostracod 4. Of the seven recognised
860 species of *Lucifer/Belzebub* (Vereshchaka et al. 2016), all have been identified in the Indo-
861 Pacific (Antony 2005; Naomi et al. 2006). Although most species seem to reach greatest
862 abundances in coastal and neritic waters rather than in the open ocean (Antony 2005), they
863 may nevertheless dominate zooplankton biomass at mesopelagic depths in oligotrophic
864 waters (Steinberg et al. 2008). That said, although Holdway and Maddock (1983) noted the
865 presence of *Lucifer* in neuston samples collected across the Indian Ocean, it was not abundant
866 – an observation in agreement here. There is limited information on the DVM of *Lucifer* in

867 oceanic waters, though it does display strong DVM in coastal waters (Woodmansee 1966;
868 Moreira et al. 2007). In the oligotrophic waters off Hawaii, Steinberg et al. (2008) noted that
869 *Lucifer* did not show pronounced DVM. In other crustacean species, the powers of DVM may
870 develop ontogenetically, with adults displaying stronger DVM than juveniles (e.g. Ringelberg
871 2009). Mature females tend to carry eggs in a loose sac for at least a short time following
872 fertilisation, and none of the specimens collected here supported any egg sacs. While it could
873 be argued that these “small” specimens were all juvenile, and so therefore perhaps less likely
874 to display pronounced DVM, it should be remembered that all species of *Lucifer/Belzebub*
875 are small in size and that the eggs sacs can be readily dislodged on capture (Naomi et al. 2006
876 and references therein).

877 Ostracods are frequently the second most abundant taxon in the plankton after copepods
878 (Angel 1999), and more than 200 species are known from different depths in the world ocean.
879 Purushothaman (2015) documents 26 species from 17 stations in the Arabian Sea, while
880 Drapun and Smith (2012) note over 40 species from more extensive sampling in the same
881 area: Graves and Angel (2017) record 41 species from a single station in the Gulf of Ohman.
882 The majority of species are found at depths greater than 200 m, but some species can be
883 collected in abundance in near surface waters (Angel 1999; Morales-Ramírez and Jakob
884 2008). Information on the DVM of planktonic Ostracods is scant, and this behaviour remains
885 undocumented for the majority of species (Angel 1999). Some species do show very
886 pronounced DVM, whilst others do not, and this behaviour does not appear to be linked to
887 occupational depth (Angel 1999). In some species, the early juvenile stages may occupy
888 shallower depths than the adults and if the species undertakes a seasonal or any other form of
889 migration, it might reflect an absence in the sampling if the full range is not covered (Angel
890 1999). Most carnivorous zooplankton will feed on Ostracods, as too will many mesopelagic
891 fishes (Uchikawa et al. 2001; Bernal et al. 2015) and the siphonophores *Hippopodius*
892 *hippopus*, *Diphyes dispur* and *Chelophyes appendiculata* appear to favour Ostracods as food
893 (Angel 1999). The higher night than day abundance of Ostracod 3 is not therefore
894 unexpected, especially as some species appear able to detect the net and to display net
895 avoidance (Angel et al. 2007), but quite why the abundance of Ostracod 4 was higher by day
896 than night is unknown. Ostracods were reported commonly in the neuston samples of
897 neuston collected by Holdway and Maddock (1983), especially from the central Indian Ocean
898 (their Area 4). Ostracod sp 3 was more abundant in neuston samples in the east of the transect
899 sampled here (Figure 8), which agrees too with the observations of Holdway and Maddock

900 (1983), who noted greatest contributions by Ostracods to the neuston in (their) Areas 2 and 3,
901 which correspond to the Banda, and Banda and Java Seas, respectively.

902 Although Rezai et al. (2019) recorded few siphonophores in the neuston of the NE Persian
903 Gulf, Holdway and Maddock (1983) found them to be common across all the transects they
904 investigated, as too they were here. Of the 23 species noted, eight were recorded in all
905 longitudinal bins (Table 6) and, as noted previously, *Abylopsis eschscholtzii* was most
906 abundant followed by *Chelophyes contorta* and *Bassia bassensis* (Table 6). Although similar
907 species (and numbers of species) were recovered amongst assemblages of very-near surface
908 siphonophores in the Colombian Pacific, Uribe-Palomino et al. (2019) noted that *Diphyes*
909 *dispar*, *Eudoxoides mitra*, *C. contorta* and *Muggiaea atlantica* were the most common
910 species, respectively. The fact that the latter of these was not recovered here reflects its
911 coastal distribution (e.g. Thibault-Botha et al. 2004), and is no surprise. Interestingly, Uribe-
912 Palomino et al. (2019) recorded few specimens of *B. bassensis*, which together with *A.*
913 *eschscholtzii* here dominated assemblages around 85° E. This species can be very abundant in
914 subtropical waters (e.g. Pagès et al. 2001) but tends to be uncommon closer to the shelf (e.g.
915 Pagès and Gill 1992). *Abylopsis eschscholtzii* dominated siphonophore assemblages in the
916 tropical Sargasso Sea (Lüskow et al. 2019), where it failed to display DVM. Neither did this
917 latter species appear to display DVM in the Celebes or Sulu Seas (Grossmann et al. 2015),
918 though these authors did note that it was one of the most shallow-living of the 90-odd species
919 recovered.

920 In their study of neuston in the North Atlantic, Hempel and Wikert (1972) observed that
921 chaetognaths made a generally small contribution to the neuston, an observation in agreement
922 with that of Brodeur et al. (1987) in the NE Pacific. Grant (1991) recorded 17 species of
923 chaetognaths in the hyponeuston of the Mid-Atlantic Bight, noting that they were generally
924 less common there than in sub-surface waters. That said, Grant (1991) did comment upon the
925 fact that species with warm-water and subtropical affinities were more regularly encountered
926 in the hyponeuston than at greater depths. By contrast, Nair et al. (1989) noted that
927 chaetognaths were common in the central Arabian Sea, Rezai et al. (2019) noted that they
928 were not uncommon in the NE Persian Gulf, and they were the second most abundant taxon
929 recorded by Holdway and Maddock (1983), especially in the coastal waters off Somalia.
930 They were the sixth most abundant taxon recovered in this study, though notably absent from
931 longitudinal bin 85, and all specimens were juveniles. This is in contrast to the results of
932 previous studies (e.g. Grant 1991), where it is assumed all collected material was adult since

933 no reference to the presence of juveniles is made. Kehayias et al. (1994) have observed that
934 juveniles of mesopelagic species of chaetognaths in the Eastern Mediterranean may occur at
935 shallower depths than adults, as too did Besiktepe and Unsal (2000) in the SW Black Sea.
936 The latter authors observed that while adult chaetognaths displayed DVM, individuals less
937 than 5 mm in length (juveniles) did not.

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943 **Table 6:** Taxa responsible for similarities within, and differences between the three clusters
 944 identified in Figure 10c.

Longitudinal Bin	95-105	Others	75-85	(95-105) vs Others		(95-105) VS (75-85)		(75-85) vs Others	
	(42,6%)	(41,75%)	(37,98%)	(61,74%)	(69,38%)	(62,91%)			
				95-105	Others	95-105	75-85	Others	75-85
Morphospecies									
<i>Abylopsis eschscholtzi</i>	0,82	1,14	1,25	0,82	1,14	0,82	1,25	0	0
Alima 1	0,54	0	0	0,54	0,38	0,54	0,09	0,38	0,09
Amphipod 1=2	0,00	0	0	0	0	0	0	0	0
Amphipod 11	0	0	0	0	0	0	0	0,17	0,17
Amphipod 12	0	0	0	0,49	0,15	0,49	0,18	0	0
Amphipod 13	0	0	0	0	0	0	0	0,13	0,2
Amphipod 16	0	0	0	0,43	0,04	0,43	0,09	0	0
Amphipod 21	0	0	0	0	0	0	0	0,23	0,08
Amphipod 9	0,68	0	0	0,68	0,56	0,68	0,19	0,56	0,19
<i>Bassia basensis</i>	0	0	0,7	0,29	0,5	0,29	0,7	0,5	0,7
Chaetognath larvae	1,27	0	0	1,27	0,97	1,27	0,38	0,97	0,38
<i>Chelophyes appendiculata</i>	0	0	0	0	0	0	0	0	0
<i>Chelophyes contorta</i>	1,05	0,91	0,91	1,05	0,78	0	0	0,78	0,91
Decapod larvae 0	0	0	0	0,48	0,69	0,48	0,18	0,69	0,18
Decapod larvae 2 (crab larvae)	0	0	0	0,55	0,19	0,55	0,08	0	0
<i>Diphyes bojani</i>	0	0	0	0,34	0,28	0,34	0,25	0,28	0,25
<i>Diphyes dispur</i>	0	0	0	0,47	0,36	0,47	0,17	0,36	0,17
Doliolid sp.	0,58	0	0	0,58	0,25	0,58	0,25	0,25	0,25
<i>Enneagonum hyalinum</i>	0	0	0	0,36	0,47	0,36	0,27	0,47	0,27
<i>Eudoxoides mitra</i>	0	0	0	0,35	0,23	0,35	0,17	0,23	0,17
<i>Eudoxoides spiralis</i>	0	0,8	0,8	0,4	0,89	0,4	0,8	0,89	0,8
Euphausiid	0	0	0	0,11	0,34	0,11	0,5	0,34	0,5
Euphausiid larvae	0	0	0	0,31	0,22	0,31	0,26	0,22	0,26
Fish B	0	0	0	0	0	0,17	0,43	0,13	0,43
Fish C	0,54	0	0	0,54	0,44	0,54	0,18	0,44	0,18
<i>Janthina</i>	0	0	0,62	0,28	0,59	0,28	0,62	0,59	0,62
<i>Lensia hotspur</i>	0	0	0	0	0	0	0	0,18	0,34
<i>Lensia subtilis</i>	0	0	0	0	0	0	0	0,19	0,17
<i>Lucifer</i>	0,75	0,75	0	0,75	0,55	0,75	0,17	0,55	0,17
Mysid	1,00	1	0	1	0,81	1	0,61	0,81	0,61
Ostracod 1	0	0	0	0	0	0,12	0,33	0,28	0,33
Ostracod 3	0,98	0,98	0,62	0,98	0,53	0,98	0,62	0,53	0,62
Ostracod 4	0	0	0	0,35	0,42	0,35	0,35	0,42	0,35
<i>Physalia</i>	0	0	0	0,45	0,09	0,45	0	0	0
<i>Porpita</i>	0	0	0	0,17	0,24	0	0	0,24	0,08
Pteropod 1	0	0	0	0	0	0	0	0	0
Pteropod 10 (Shell A)	0	0,5	0	0,5	0,59	0,5	0,18	0,59	0,18
Pteropod 11	0	0	0	0,51	0,43	0,51	0,63	0,43	0,63
Pteropod 3	0	0	0	0	0	0	0	0	0
Pteropod 8 (Atlanta sp)	0	0	0	0,35	0,27	0	0	0,27	0,17
<i>Sulculiolaria biloba</i>	0	0	0	0,29	0,17	0,29	0,25	0,17	0,25
<i>Thalia</i>	0,70	0	0	0,7	0,75	0,7	0,46	0,75	0,46
unidentified fish larvae	0,86	0,86	0	0,86	0,62	0,86	0,17	0,62	0,17
<i>Vellela</i>	0	0	1,12	0,42	1,03	0,42	1,12	0	0
Water strider A	0,53	0,53	0	0,53	0,19	0,53	0,08	0	0
Water strider B	0,88	0	0	0,88	0,39	0,88	0,17	0,39	0,17
Water strider C	0,56	0	0	0,56	0,14	0,56	0,17	0	0
Zooanthidae larvae	0	0	0	0,33	0,23	0,33	0,08	0,23	0,08

945

946 Meroplankton was represented by a wide variety of taxa but was dominated, for the most
947 part, by larvae of Decapods, Stomatopods and fish. Larvae of brachiopods, polychaetes,
948 bryozoans, gastropods, bivalves and echinoderms were rare, though they were conspicuous in
949 the study of Holdway and Maddock (1983) from across this same area. Indeed, the latter
950 authors noted that meroplankton represented about 33% of the non-copepod neuston, whereas
951 here they constituted between 9% and 21%, being more abundant at the longitudinal extremes
952 of the transect than at the centre. If we accept that waters along the transect have their origin,
953 in part, from the east and from the Indonesian throughflow jet (Dufois et al. 2017 and
954 references therein), then perhaps the high numbers of meroplankton found in samples
955 collected close to Jakarta and Christmas Island can be explained by their proximity to the
956 coral or East Indies triangle (Hoeksema 2007; Veron et al. 2009). Coral reefs support very
957 rich associated communities, especially of Decapods (Klomp maker et al. 2013 and references
958 therein) and Stomatopods (Barber et al. 2002), many of which have planktonic larval
959 development. As this advected water moves westward from the coast and into open ocean
960 water, it is likely that the meroplanktonic species drop out of assemblages, as has been noted
961 by Ayata et al. (2011) for polychaetes in the Bay of Biscay and Meerhof et al. (2018) for
962 Decapod larvae around Easter Island. It is then perhaps no surprise that the number of such
963 benthic meroplankton are low between 75-85° E. Moving further westward still, the numbers
964 of zoanthid and Decapod larvae and Stomatopod alima once again begin to increase, which
965 could reflect the southward advection of waters from the Mascarene Plateau, because the
966 diversity of both taxa is known to be relatively high off the east coast of Africa (Reaka et al.
967 2008; Head et al. 2018).

968 Mysids were found across the sampling area, being slightly more common in the east than the
969 west, which is subtly at odds with the observations of Holdway and Maddock (1983). The
970 latter authors noted that these crustaceans were more abundant in the Indian Ocean than
971 elsewhere, but that they were slightly more common close to the African coastline than they
972 were in the Banda and Java Seas. Mysids were significantly more common in samples
973 collected at night than they were during the day, which agrees with previous observations
974 (e.g. Hempel and Wieckert 1972; Holdway and Maddock 1983). While mesopelagic species
975 of mysid may not display pronounced DVM (Andersen and Sardou 1992; Hopkins et al.
976 1994), this behaviour is frequently noted by shallow water (e.g. Kouassi et al. 2006) and lotic
977 species (Beeton and Bowers 1982). Epipelagic species of mysid are generally uncommon in
978 the open ocean and are characterised by *Siriella thompsoni* and *Anchialina typica* (Murano

979 and Mauchline 1999), which tend to be associated with warm water. Hempel and Wieckert
980 (1972) noted that both species descended to about 200 m during the day before reappearing in
981 surface waters at night.

982 Whilst doliolids were scarce, salps, including morphospecies of *Thalia* (as well as
983 *Traustedia* and *Salpa*) were relatively common across the sampling area, and were slightly
984 more numerous in night-time than day-time tows. Holdway and Maddock (1983) also noted
985 the presence of both taxa in their Indian Ocean samples, at frequencies that what would
986 appear to be similar to those noted here. That said, few other students of neuston have
987 remarked on the presence of thaliaceans in their collections, despite the fact that these
988 organisms are known to graze microplastics very effectively (Brandon et al. 2020). Although
989 some species of salp are known to display pronounced DVM for reproductive purposes
990 (Madin et al. 1996), it is clear that others do not (Gibbons 1997) and attempts to interpret the
991 data recovered here would be speculative in this regard. More work on this group remains to
992 be done, because they are not generally regarded as members of the neuston (but see also
993 Mantha et al. 2019).

994 Tropical and warm water assemblages of pelagic amphipods are typically more diverse than
995 those in cooler waters (Vinogradov 1999), and after Copepoda, Amphipoda is probably the
996 most diverse group of holopelagic crustaceans. Given this, their richness in neustonic samples
997 is perhaps no surprise. The absence of any distinct latitudinal pattern in amphipod assemblage
998 structure across the transect agrees with the observations of Vinogradov (1999). The latter
999 author noted that "... species compositions are relatively homogeneous within the vast
1000 tropical anticyclonic gyres, and dominant species are also more or less similar in the
1001 various collections...". The distinct diel difference in assemblages undoubtedly reflects
1002 DVM (Vinogradov 1999; Holdway and Maddock 1983) and the movement into the surface
1003 waters of species that typically reside at greater depths during the day.

1004 The sea-surface represents an important nursery area for many epi- and mesopelagic fish
1005 species, and a wide variety of forms can be found there, usually as eggs and larvae (e.g.
1006 Doyle 1992), where they may represent important diet items for marine birds, mammals and
1007 larger fishes (Olivar et al. 2015). Although there are relatively few studies of the neustonic
1008 ichthyofauna in the open Indian Ocean, Olivar et al. (2015) recorded 14 larval taxa in their
1009 transect from west to east at ~25°S, and assemblages were dominated by Myctophidae and
1010 Exocoetidae. Here, only eleven morphospecies were recovered (no eggs), in very low

1011 numbers, and although most failed to display clear day-night differences in abundance (as
1012 Holdway and Maddock 1983, but unlike Olivar et al. 2015), Morphospecies C was more
1013 common by day than night. In the absence of any identification, however, it is not possible to
1014 explain this behaviour. Neustonic fish larvae are notoriously patchy in their distribution
1015 (Olivar et al. 2015), and this could explain the absence of any spatial pattern to the
1016 distribution of assemblages here. That said, diversity and abundance were greater at the
1017 eastern part of the transect than in the centre or to the west, probably as a result of
1018 contributions from the Indonesian Throughflow jet (see above).

1019 Hempel and Weikert (1972) also observed species of *Creseis* to be more common by day
1020 than night, while Holdway and Maddock (1983) noted heteropods to be more more common
1021 at night than day. Diel vertical migration by Pteropods was noted by van der Spoel and
1022 Dadon (1999), while Richter and Seapy (1999) refer to the occasional abundance of
1023 heteropods in neuston samples, which is thought to be for reproductive purposes.
1024 Interestingly, the latter authors report that *Halobates* may use floating heteropod egg masses
1025 for oviposition. Sea-swallows (*Glaucus* sp.) were rare, and while *Janthina* were recorded
1026 across the region they were slightly more common to the west. Both these latter species are
1027 predators of the neustonic hydrozoans *Porpita*, *Velella* and *Physalia* (Lepoint et al. 2016;
1028 Pinotti et al. 2019), but there was no correlation (Pearson's *R*) between the abundance of the
1029 two groups ($p>0.05$).

1030

1031 Less common taxa

1032 Isopods are known to occur in neuston samples, though not frequently and not in large
1033 numbers (Holdway and Maddock 1983; Brodeur et al. 1987). As here, where a total of 11
1034 individuals from four morphospecies were observed (Table 1; Figure 13). Isopods of the
1035 genus *Idotea* and *Eurydice* are typically more common in neuston samples from near-coastal
1036 waters (Holdway and Maddock 983; Brodeur et al. 1987; Echelman and Fishelson 1990),
1037 where they can also be associated with rafting algae and other flotsom (Thiel and Gutow
1038 2004 and references therein). The absence of *Sargassum* or other floating debris may account
1039 for the general scarcity of isopods in our samples.

1040 Euphausiids were relatively uncommon across the transect (Table 1) and only two
1041 morphospecies were identified, putatively *Euphausia brevis* and *Stylocheiron* sp. While

1042 neither morphospecies was caught regularly in the centre of the transect, the former was more
1043 common in the west and the latter in the east. Both species were slightly more abundant by
1044 night than day, an observation in agreement with previous studies (Holdway and Maddock
1045 1983; Brodeur et al. 1987; Echelman and Fishelson 1990). As in previous studies, larvae
1046 (some calyptopes but mostly furcilia) were always more abundant than adults, but like adults
1047 they too were more common at night than by day (as Holdway and Maddock, 1983).
1048 Euphausiids are often regarded as micronekton and are known to perform pronounced DVM
1049 (e.g. Pillar et al. 1992).

1050 Although polychaetes occurred at low average abundances, five morphospecies were
1051 observed. These included one *Tomopteris*, three Alciopidae and miscellaneous larvae
1052 (considered here as one morphospecies) (Table 1: Figure 13). With the exception of the
1053 larvae, about which little can be said, the other taxa are regarded as holoplanktonic
1054 (Fernández-Alamo and Thuesen 1999). The absence of (e.g.) nereids suggests floating
1055 substrata were absent from neuston samples (as noted too above with regard to isopods), as
1056 these benthic taxa may commonly occur amongst *Sargassum* and other flotsom (Baker et al.
1057 2018). The taxa recorded here were noted too by Brodeur et al. (1987). Polychaetes were
1058 more common towards the east (Holdway and Maddock 1983) and were found exclusively at
1059 night (Table 7). Information regarding the DVM of pelagic polychaetes is thin: deeper-living
1060 species of *Tomopteris* in cooler water do not appear to display any change in vertical
1061 distribution (Guglielmo et al. 2014), though tomopterids and alciopids in the Sargasso Sea
1062 have been shown to be very strong migrators (Eden et al. 2009).

1063 Gastropod larvae (of unknown origin) were uncommon (Figure 13), as observed too by
1064 Holdway and Maddock (1983) in the Indian Ocean, which is perhaps no surprise if they are
1065 the product of benthic forms. In neuston samples collected in coastal waters of the northern
1066 Red Sea and Gulf of Aqaba, however, Gastropoda (identity not stated) were almost as
1067 abundant as copepods, and were more abundant at night than by day (Mantha et al. 2019).
1068 Given that bivalve larvae were also very numerous in the latter study, as too were a number
1069 of other meroplanktonic forms, it is likely that the Gastropoda recovered by Mantha et al.
1070 (2019) were larvae. Liparoto et al. (2017) found in the Western Mediterranean and Southern
1071 Adriatic Sea that apart from the eueneuston, gastropod veligers were the most widely
1072 distributed taxa and were found at all the stations sampled. Which again is no surprise given
1073 the neritic nature of the environment sampled.

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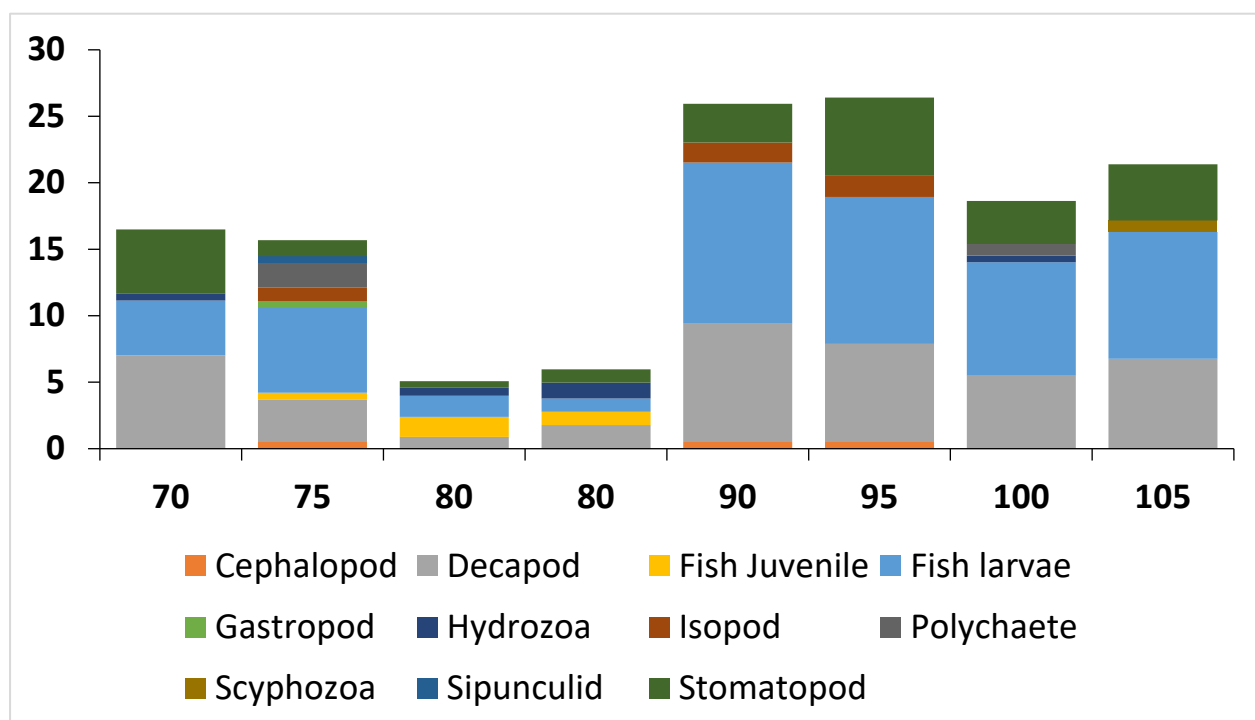


Figure 13: Average abundance of non-major taxa across all longitudinal classes.

1083
1084 **Diel vertical migration**

1085 The diversity and abundance of the neuston was higher at night than during the day, as
1086 euneuston was supplemented with facultative neuston following DVM (Figures 6a, c). As
1087 noted above in reference to the individual taxa, many facultative members of the neuston
1088 move into the surface waters at night in order to feed and escape predators (Williamson et al.
1089 1996). The significant role that the time of day plays in shaping the structure of neuston
1090 assemblages has long been recognised (Hempel and Weikert 1972), but it means that unless
1091 there is an equal balance of samples across day and night time-series, it is very difficult to
1092 rigorously investigate the role played by any other factors simultaneously. Hence, as here,
1093 further analyses were partitioned by time of day.

1094 The marginal tests of the DistLM indicate aspects of the chlorophyll environment (sea
1095 surface fluorescence, integrated fluorescence, depth of the fluorescence maximum and value
1096 of fluorescence at the fluorescence maximum) were all correlated with assemblage structure
1097 at night. But not during the day. The euneuston is dominated by carnivores: *Physalia* (Bieri
1098 1970; Holdway and Maddock 1983), *Velella* (Purcell et al. 2015; Betti et al. 2019), *Porpita*
1099 (Sahu et al. 2020), *Janthina* (Churchill et al. 2011), *Glaucus* (Thompson and Bennett 1970;
1100 Sahu et al. 2020), *Halobates* (Cheng 1985), as are chaetognaths (Stuart and Verheye 1991;
1101 Albuquerque et al. 2021), siphonophores (Mackie et al. 1988) and most hyperiid amphipods

1102 (Shulenberg 1977). The link between these species and the fluorescent environment is
1103 therefore indirect. By contrast, herbivorous/detritivorous taxa (including Pteropods, salps,
1104 mysids, euphausiids and Ostracods), which tended to be at least twice as common in neuston
1105 samples at night than by day (Table 7), are likely to be directly influenced by fluorescence. It
1106 is perhaps not surprising then that overall assemblage structure was more influenced by
1107 fluorescence at night than during the day. That said, in the full DistLMs generated for night
1108 and day assemblages, aspects of the fluorescence environment were incorporated into both
1109 (with similar overall weightings, Tables 4b and 5b), implying perhaps that its impact is “felt”
1110 both directly and indirectly.

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1112



1113 **Table 7:** Diel changes in the abundance of major taxa, averaged across samples per station;
 1114 data as root-root numbers per trawl.

Taxon	Day	Night
Amphipod	1,232	4,201
Barnacle	0	0
Cephalopod	0	0,077
Chaetognath	0,983	0,955
Decapod	1,368	1,668
Doliolid	0,271	0,349
Euphausiid	0,296	0,787
Fish Juvenile	0,067	0,051
Fish larvae	1,479	1,328
Gastropod	1,871	3,025
Heteropod	0,033	0
Hydrozoa	0,980	1,141
Isopod	0,067	0,155
Mysid	0,478	1,026
Ostracod	0,929	1,713
Polychaete	0,000	0,316
Salp	0,582	1,191
Scyphozoa	0	0,107
Siphonophore	5,509	6,796
Sipunculid	0	0,026
Stomatopod	0,209	0,813
Water strider	0,590	1,162
Zooanthidae	0,170	0,205



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1118 Longitudinal patterns displayed by neuston
1119 The use of morphospecies complicates clear interpretation of the longitudinal patterns
1120 because, as noted earlier, a single morphospecies may comprise more than one sibling
1121 species, whilst different developmental stages of the same species, may be regarded here as
1122 representing separate morphospecies. The latter is especially likely to occur in the case of
1123 larval Decapods, polychaetes and even fish larvae, whose anatomy and outward appearance
1124 may change quite markedly during development (Williamson 2013). The other taxon that is
1125 particularly problematic in this regard is Pteropoda, species of which can usually be identified
1126 by the nature of their calcareous shell (e.g. Boltovskoy et al. 1996 cited in Boltovskoy 1999).
1127 Unfortunately, following prolonged exposure to formalin, the shells of Pteropods had all but
1128 dissolved, and animals were separated on the basis of the shape of the remaining body mass
1129 and wing-like feet; when the latter had not been retracted. Regardless, every effort was made
1130 to be consistent in specimen categorisation to morphospecies and so it is hoped that errors
1131 were consistent across the samples.

1132 If we accept the above, and are conservative in our interpretation of the results, then the
1133 results suggest the following. Firstly, that assemblages between 75 - 85° E (corresponding to
1134 Grids 80 and 85) are distinct (Figures 9 and 10). Secondly, that the easternmost assemblages
1135 of the balance (corresponding to Grids 100 and 105) are distinct (Figure 9c), and thirdly that
1136 there is an approximate symmetry of the remaining samples around 75 - 85° E (Figure 9c).

1137 During the 12 months prior to the survey, the “centre” of the Southern Indian Ocean Gyre lay
1138 approximately along latitude 20°S, and extended from ~95° E westwards to the coastline of
1139 Madagascar at ~50° E. At the time of the cruise, its mean position was strongly disrupted by
1140 a series of mesoscale eddies (both cyclonic and anticyclonic) ranging in size from 50 - 150
1141 km diameter (Torens et al. 2015). Two distinct, anticyclonic, warm core eddies, with
1142 elevated sea surface heights were detected between 75 - 85° E (Dufois et al. 2017).

1143 Conditions in these eddies were distinct from those at the edge, with elevated concentrations
1144 of chlorophyll being attributed to deeper convective mixing (Dufois et al. 2017).

1145 Assemblages of plankton in the centre of eddies are often distinct from those outside (Dufois
1146 et al. 2017; Bernal et al. 2020) due to the horizontal advection of productive waters and
1147 deeper vertical mixing in the anticyclonic eddies causing an increase in surface chlorophyll in
1148 the anticyclonic eddies in winter in the Southern Indian Ocean Gyre (Dufois et al. 2017). And
1149 it is likely that the neuston were behaving similarly here. Interestingly, however, assemblages
1150 in this region were generally less abundant, species-rich and diverse than those to either the

1151 east or the west, and were characterised by relatively high numbers of siphonophores and
1152 some holozooplankton (Ostracod 1, Pteropod 11) but low levels of all meroplankton except
1153 fish larva 4. The near absence of meroplankton is perhaps no surprise, given that the majority
1154 of taxa would likely be of benthic forms (Schnack-Schiel and Isla 2005) and the great
1155 distance of this region from the coast; as has been argued above.

1156

1157 The high diversity and distinct nature of assemblages between 95 - 105° E no doubt reflects
1158 the impact of water emerging from the Java and Banda Seas in the Indonesian throughflow.
1159 As noted previously, the East Indies Triangle is a global biodiversity hotspot (e.g. Roberts et
1160 al. 2002) for many benthic or reef associated taxa and is likely to have levels of plankton
1161 diversity too (e.g. Yasuhara et al. 2012). This area was characterised by relatively high
1162 temperatures and low salinities at the surface and neuston assemblages delimited by water-
1163 striders, doliolids and high numbers of chaetognath larvae, *Lucifer* and mysids: zoanthid
1164 larvae were also common. Morphospecies of *Thalia* were also common, but like Pteropod 10,
1165 they were also common in the longitudinal bins around the core at 75 - 85° E.

1166

1167



1168 **Conclusions**

1169

1170 Overall, the results of the study were in accordance with much of the previous literature in
1171 terms of the general taxa present, though they differed in some respects. The preponderance
1172 of Ostracods and juvenile chaetognaths in samples was unusual, while the lack of flotsam
1173 meant that many of those organisms that inhabit such structures were missing. The fact that
1174 assemblages were generally more diverse towards the margins of the transect is no surprise,
1175 given that holoplankton was being supplemented with meroplankton – especially in the east
1176 with water from the Indonesian Throughflow: this is as originally hypothesised. Communities
1177 were strongly influenced by the time of day, reflecting the diel vertical migration of
1178 facultative members, whose abundances and diversity were all greater at night: which again
1179 comes as no surprise.

1180 There was a strong agreement between the structure of assemblages and the structure of the
1181 physical environment, and it is clear that assemblages in the core of the anticyclonic eddies
1182 were distinctly different from those outside, as initially hypothesised. In part this may reflect
1183 the drop-out of meroplanktonic forms, but there is a suggestion in the data that the
1184 chlorophyll environment may have some role to play in this. More work that couples the
1185 physical environment with the neuston is clearly needed, though this study emphasises the
1186 value of simultaneously collected data sets in this regard.

1187 This study used morphospecies as identification units, rather than recognised species (with the
1188 exception of siphonophores), and I believe this approach holds much promise for future
1189 work. It does have its drawbacks as articulated previously, but it does allow, by comparison
1190 with (e.g.) ZooScan, a greater resolution in taxonomic identification (Gorsky et al. 2010),
1191 without the need for full-blown training in the taxonomy of individual taxa. It is unlikely to
1192 be as thorough in its resolution of species units as e-DNA (Garcia-Vasquez et al. 2021) or
1193 molecular approaches to identification (Bucklin et al. 2021), but in an African context where
1194 the cost of such technologies is so high (let alone the specialist laboratories needed to
1195 undertake the work), it does have its advantages (see also Machida et al. 2021).

1196

1197 **References**

1198

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1589 Appendix A

1590 List of morphospecies recovered from neuston samples collected during the IOS Leg 1
 1591 survey of the southern Indian Ocean, June-July 2015. Also shown are higher levels of
 1592 taxonomic classification, the number of CTD stations from which each was recovered and
 1593 their allocated numbers (see Appendix B).

1594

Number	Morphospecies	Taxon 1	Holo/Mero	Epi/Hypo/Fac	No stations
1	Amphipod 10	Amphipod	Holoplankton	facultative	8
2	Amphipod 11	Amphipod	Holoplankton	facultative	12
3	Amphipod 12	Amphipod	Holoplankton	facultative	11
4	Amphipod 13	Amphipod	Holoplankton	facultative	11
5	Amphipod 15	Amphipod	Holoplankton	facultative	2
6	Amphipod 16	Amphipod	Holoplankton	facultative	6
7	Amphipod 17	Amphipod	Holoplankton	facultative	5
8	Amphipod 18	Amphipod	Holoplankton	facultative	9
9	Amphipod 19	Amphipod	Holoplankton	facultative	3
10	Amphipod 20	Amphipod	Holoplankton	facultative	6
11	Amphipod 21	Amphipod	Holoplankton	facultative	14
12	Amphipod 22	Amphipod	Holoplankton	facultative	1
13	Amphipod 23	Amphipod	Holoplankton	facultative	2
14	Amphipod 24	Amphipod	Holoplankton	facultative	2
15	Amphipod 25	Amphipod	Holoplankton	facultative	1
16	Amphipod 26	Amphipod	Holoplankton	facultative	2
17	Amphipod 27	Amphipod	Holoplankton	facultative	1
18	Amphipod 6	Amphipod	Holoplankton	facultative	10
19	Amphipod 7	Amphipod	Holoplankton	facultative	3
20	Amphipod 8	Amphipod	Holoplankton	facultative	4
21	Amphipod 9	Amphipod	Holoplankton	facultative	20
22	<i>Phronima</i> Amphipod 5	Amphipod	Holoplankton	facultative	4
23	<i>Synopia</i> Amphipod 1=2	Amphipod	Holoplankton	facultative	21
24	<i>Vibilia</i> Amphipod 4	Amphipod	Holoplankton	facultative	6
25	Barnacle larvae - cypris	Barnacle	Meroplankton	facultative	1
26	Cephalopod larvae 1	Cephalopod	Meroplankton	facultative	2
27	Cephalopod larvae 2	Cephalopod	Meroplankton	facultative	2
28	<i>Spirula</i>	Cephalopod	Holoplankton	Hypo	1
29	Chaetognath larvae	Chaetognath	Holoplankton	facultative	26
30	Sergestidae	Decapod	Holoplankton	facultative	4
31	Decapod larvae 0	Decapod	Meroplankton	facultative	23
32	Decapod larvae 10	Decapod	Meroplankton	facultative	1
33	Decapod larvae 12	Decapod	Meroplankton	facultative	1
34	Decapod larvae 2	Decapod	Meroplankton	facultative	11
35	Decapod larvae 5=6	Decapod	Meroplankton	facultative	4
36	<i>Lucifer</i>	Decapod	Holoplankton	facultative	20
37	Doliolid sp.	Doliolid	Holoplankton	facultative	17
38	<i>Euphausia brevis</i>	Euphausiid	Holoplankton	facultative	1
39	Euphausiid	Euphausiid	Holoplankton	facultative	14

40	Euphausiid larvae	Euphausiid	Holoplankton	facultative	15
41	<i>Stylocheiron</i> (furcilia)	Euphausiid	Holoplankton	facultative	3
42	Fish A	Fish Juvenile	Meroplankton	facultative	1
43	Fish I	Fish juvenile	Meroplankton	facultative	2
44	Fish J	Fish juvenile	Meroplankton	facultative	1
45	Fish K	Fish juvenile	Meroplankton	facultative	1
46	Fish B	Fish larvae	Meroplankton	facultative	8
47	Fish C	Fish larvae	Meroplankton	facultative	19
48	Fish D	Fish larvae	Meroplankton	facultative	3
49	Fish E	Fish larvae	Meroplankton	facultative	2
50	Fish F	Fish larvae	Meroplankton	facultative	1
51	Fish G	Fish larvae	Meroplankton	facultative	6
52	Fish H	Fish larvae	Meroplankton	facultative	4
53	Fish I	Fish larvae	Meroplankton	facultative	22
54	<i>Creseis</i> sp (1)	Gastropod	Holoplankton	facultative	9
55	<i>Creseis</i> sp (2)	Gastropod	Holoplankton	facultative	2
56	<i>Creseis</i> sp (3)	Gastropod	Holoplankton	facultative	13
57	Gastropod larvae	Gastropod	Meroplankton	facultative	1
58	<i>Glaucus</i> sp.	Gastropod	Holoplankton	Hypo	1
59	<i>Janthina</i>	Gastropod	Holoplankton	Hypo	20
60	LargePteropod 15	Gastropod	Holoplankton	facultative	1
61	Pteropod 1	Gastropod	Holoplankton	facultative	7
62	Pteropod 10 (Shell A)	Gastropod	Holoplankton	facultative	19
63	Pteropod 11 (Shell B)	Gastropod	Holoplankton	facultative	19
64	Pteropod 12	Gastropod	Holoplankton	facultative	2
65	Pteropod 13	Gastropod	Holoplankton	facultative	1
66	Pteropod 14	Gastropod	Holoplankton	facultative	1
67	Pteropod 2	Gastropod	Holoplankton	facultative	4
68	Pteropod 4	Gastropod	Holoplankton	facultative	3
69	Pteropod 5	Gastropod	Holoplankton	facultative	2
70	Pteropod 6	Gastropod	Holoplankton	facultative	3
71	Pteropod 7	Gastropod	Holoplankton	facultative	1
72	Pteropod 8 (<i>Atlanta</i> sp)	Gastropod	Holoplankton	facultative	15
73	Pteropod 9	Gastropod	Holoplankton	facultative	8
74	Heteropod larvae (pterotrachidae)	Heteropod	Holoplankton	facultative	1
75	<i>Bougainvillia</i> sp	Hydrozoa	Meroplankton	facultative	4
76	<i>Porpita</i>	Hydrozoa	Holoplankton	epi	11
77	Rhopalonematidae	Hydrozoa	Holoplankton	facultative	1
78	<i>Vellela</i>	Hydrozoa	Holoplankton	epi	25
79	Isopod	Isopod	Meroplankton	facultative	8
80	Mysid	Mysid	Holoplankton	facultative	24
81	Ostracod 1	Ostracod	Holoplankton	facultative	9
82	Ostracod 3	Ostracod	Holoplankton	facultative	19
83	Ostracod 4	Ostracod	Holoplankton	facultative	18
84	Ostracod 5	Ostracod	Holoplankton	facultative	5
85	Ostracod 7	Ostracod	Holoplankton	facultative	2
86	<i>Vanadis</i> sp2)	Polychaete	Holoplankton	facultative	3
87	Polychaete 2	Polychaete	Holoplankton	facultative	2
88	Polychaete 3 (larvae)	Polychaete	Meroplankton	facultative	2

89	<i>Vanadis</i> sp1	Polychaete	Holoplankton	facultative	2
90	<i>Tomopteris</i> sp	Polychaete	Holoplankton	facultative	1
91	<i>Iasis zonaria</i>	Salp	Holoplankton	facultative	5
92	<i>Pegea</i>	Salp	Holoplankton	facultative	2
93	<i>Salpa</i>	Salp	Holoplankton	facultative	7
94	<i>Thalia</i>	Salp	Holoplankton	facultative	27
95	<i>Trastedia multitentaculata</i>	Salp	Holoplankton	facultative	2
96	<i>Pyrosoma</i>	Salp	Holoplankton	facultative	2
97	Ephyrae	Scyphozoa	Meroplankton	facultative	1
98	<i>Nausithoe</i> sp	Scyphozoa	Holoplankton	facultative	1
99	<i>Pelagia</i> sp.	Scyphozoa	Holoplankton	facultative	3
100	Scyphozoan 1	Scyphozoa	Holoplankton	facultative	1
101	Scyphozoan 2	Scyphozoa	Holoplankton	facultative	1
102	<i>Abyla trigona</i>	Siphonophore	Holoplankton	facultative	6
103	<i>Abylopsis eschscholtzi</i>	Siphonophore	Holoplankton	facultative	27
104	<i>Abylopsis tetragona</i>	Siphonophore	Holoplankton	facultative	9
105	<i>Agalma elegans</i>	Siphonophore	Holoplankton	facultative	1
106	<i>Agalma okeni</i>	Siphonophore	Holoplankton	facultative	1
107	<i>Bassia basensis</i>	Siphonophore	Holoplankton	facultative	18
108	<i>Chelophyes appendiculata</i>	Siphonophore	Holoplankton	facultative	17
109	<i>Chelophyes contorta</i>	Siphonophore	Holoplankton	facultative	27
110	<i>Diphyes bojani</i>	Siphonophore	Holoplankton	facultative	17
111	<i>Diphyes dispur</i>	Siphonophore	Holoplankton	facultative	18
112	<i>Enneagonum hyalinum</i>	Siphonophore	Holoplankton	facultative	22
113	<i>Erenna</i> sp.	Siphonophore	Holoplankton	facultative	1
114	<i>Eudoxoides mitra</i>	Siphonophore	Holoplankton	facultative	11
115	<i>Eudoxoides spiralis</i>	Siphonophore	Holoplankton	facultative	26
116	<i>Hippopodius hippopus</i>	Siphonophore	Holoplankton	facultative	6
117	<i>Lensia campanella</i>	Siphonophore	Holoplankton	facultative	2
118	<i>Lensia hotspur</i>	Siphonophore	Holoplankton	facultative	7
119	<i>Lensia subtilis</i>	Siphonophore	Holoplankton	facultative	6
120	<i>Physalia</i>	Siphonophore	Holoplankton	epi	9
121	<i>Sulculiolaria biloba</i>	Siphonophore	Holoplankton	facultative	12
122	<i>Sulculiolaria chuni</i>	Siphonophore	Holoplankton	facultative	8
123	<i>Vogtia glabra</i>	Siphonophore	Holoplankton	facultative	4
124	Sipunculid larvae	Sipunculid	Meroplankton	facultative	1
125	Alima 1	Stomatopod	Meroplankton	facultative	17
126	Alima 2	Stomatopod	Meroplankton	facultative	1
127	Alima 3	Stomatopod	Meroplankton	facultative	3
128	Stomatopod 1	Stomatopod	Meroplankton	facultative	2
129	Stomatopod 2	Stomatopod	Meroplankton	facultative	2
130	Stomatopod 3	Stomatopod	Meroplankton	facultative	1
131	Stomatopod 4	Stomatopod	Meroplankton	facultative	3
132	Water strider A	Water strider	Holoplankton	epi	12
133	Water strider B	Water strider	Holoplankton	epi	20
134	Water strider C	Water strider	Holoplankton	epi	12
135	Zooanthidae larvae	Zooanthidae	Meroplankton	facultative	14



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1597 Appendix B.
1598 Average abundance (root-root numbers) of neuston morphospecies (see Appendix A for
1599 details) recovered from samples collected during the IOS Leg 1 survey of the southern Indian
1600 Ocean, June-July 2015. Also shown are details regarding the time of day (day, D; Night, N;
1601 Dusk, Dk; Dawn, Dn) of sample capture and gridded longitudinal bin (e.g. 65 - 70° E = 70).
1602



Longitude Bin	105	105	105	100	100	100	95	95	95	95	95	90	90	90	90	90	85	85	80	80	75	75	75	75	75	70	70	
Time of Day	D	N	N	N	D	Dk	D	Dk	N	D	N	N	D	D	N	N	Dk	N	N	D	D	N	D	Dk	N	Dn	N	
Number	500	503	504	506	508	509	511	512	513	514	515	516	517	518	519	520	523	524	527	529	531	532	534	535	536	537	539	
1	0,439	0	0	0	0	0,498	0	0	0,396	0	0	0	0	0	1,016	0	0	0,333	0	0	0	1,917	0	0,978	0	0	0	
2	0	0	0,333	0,835	0	0	0	0,333	0,333	0	0,396	0,333	0,439	0	0,333	0	0,333	0	0,333	0	0	0,333	0	0	0	0	1,029	
3	0	1,217	0,333	1,593	0,333	0,498	0	2,349	0	0	0	0	0	0	0	0	0,396	0	0,522	0	0,333	0	0	0	0,333	0	0,895	
4	0	0,73	0	0,439	0	0,835	0	0	0	0,333	0	0,396	0	0	0,333	1,351	0	0	0	0	0	0	0	0,333	0,333	0	0,439	
5	0	0	0	0	0	0,333	0	0	0	0	0	0	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	
6	0	1,84	0	0,855	0	0,73	0	0	0	0,333	0	0	0	0	0	0,439	0	0	0	0	0	0	0	0	0,333	0	0	
7	0	0,439	0	0,396	0	1,083	0	0,911	0	0	0	0	0	0	0	0	1,503	0	0	0	0	0	0	0	0	0	0	
8	0	0,498	0	0,333	0	0,333	0	1,107	0	0	0,439	0	0	0,667	0	0	0,333	0	0	0	0	0	0	0	0	0,593	0	0,439
9	0	0	0	0	0	0	0	0,333	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,333	0	0
10	0,333	0	0	0,333	0	0,835	0	0,894	0	0	0	0	0	0,333	0	0	0,561	0	0	0	0	0	0	0	0	0	0	
11	0,838	0,333	0	0,333	0	0	0,471	0,333	0,396	0	0,667	0,333	0	0	0,333	0	0,333	0	0	0	0	0,333	0	0	0	0,333	0,471	
12	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,333	0	0	
13	0	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,333	0	0	
14	0	0	0	0	0	0,561	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
15	0	0	0	0	0	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
16	0	0	0	0	0	0,705	0	0	0	0	0	0	0,607	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
17	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,333	
18	0	0,498	0	0,396	0	0	0	1,12	0	0	0,498	0	0	0	0	0	0	0	0,333	0	0	0,396	0	0	0,939	0,439	0,793	
19	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,396	0	0	0	0,396	0	0	0,333	0	0	0	0	0	
20	0,333	0	0	0	0	0	0	0	0,333	0	0	0	0	0	0,396	0	0	0	0	0	0	0,333	0	0	0	0	0	
21	1,357	1,679	0,333	1,417	0	0,805	0	0,471	1,833	0,333	0,772	0,835	0	0	1,534	0,667	0,471	0	0,561	0	0,333	1,165	0	1,307	1,334	0	1,184	
22	0	0	0	0	0	0,73	0	0	0	0	0	0	0,333	0	0	0	0	0	0,333	0	0	0	0	0	0,333	0	0	
23	0	2	0	3,083	0,835	2,905	0	0,333	2,615	0	2,251	1,29	1,636	0,396	2,309	1,876	1,249	0,667	1,863	0	0	1,742	0,333	1,59	2,09	0	2,516	
24	0	0	0	0,667	0	0	0	0	0,333	0	0	0	0	0	0,396	0	0	0,396	0	0	0	0,333	0	0	0	0	0,396	
25	0	0	0	0	0	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
26	0	0	0	0	0	0	0	0,333	0	0	0	0	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	
27	0	0	0	0	0	0	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0,333	0	0	0	0	0	
28	0	0	0	0	0	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
29	3,193	2,987	1,738	3,411	2,648	1,959	2,991	4,226	3,317	1,594	2,228	1,61	1,333	2,258	2,047	0,868	0,333	0	2,014	0	0,439	0,396	0,793	0,753	0,396	1,251	1,564	
30	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,333	0	0,333	0	0	0,993	0	0	0	0	0,396	
31	0,793	1,839	0	0,73	0,396	0	0,73	1,475	1,719	0,73	0,832	1,291	0,73	1,059	0,73	0,333	0	0,333	0,593	0	0,333	0	0,333	0,772	0,667	0,333	1,063	
32	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
33	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,396	0	0	0	0	0	0	0	0	0	0	0	0	
34	0,333	1,546	0	1,988	0	0,772	0	0	0,333	0	0	0	0,333	0	0	0,73	0,333	0	0	0	0	0	0	0	0	0,805	0,333	1,461
35	0	0	0	0,542	0	0	0	1,66	0	0,577	0	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	
36	3,035	0,333	0,832	1,452	0	1,231	1,89	2,072	0,957	1	0,396	0,667	0	1,052	0,333	0,333	0	0	0,805	0	0,868	0	0	0	0,333	0,333	0,835	
37	0,793	1,455	0	0,772	0,333	0,73	0,667	1,2	0,396	0	0	0,396	0	0,333	0	0,333	0,396	0,333	0	0,333	0	0	0	0	0,333	0	1,274	
38	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,498	0	0	0	0	0	0	0	0	0	0	0	
39	0	0	0	0	0	0,73	0	0,333	0,333	0,333	0,333	0	1,487	0	0,542	0,667	1,736	0,396	1,485	0	0	0	0	1,082	1,219	0	1,825	
40	1,038	0,333	0	0,822	0	0,439	0,333	0,471	0,667	0	0,333	0,439	0	0	0,471	0	0,439	0	0,73	0	0	0	0	0,805	0,439	0	1,386	
41	1,157	0	0	0	0	0	0	0	0,772	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,333	
42	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,333	0	0	0	0	0	0	0	0	
43	0	0	0	0	0	0,333	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
44	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,667	0	0	0	0	0	0	0	
45	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,333	0	0	
46	0	0	0	0	0	1,063	0	0	0	0	0	0,667	0,333	0	1,063	0	1,201	0,333	0	0,333	0	0,333	0	0	0	0	0	
47	0	0,333	0,895	0,396	2,004	0,73	0	0,333	0,333	0,667	0,73	0	1,211	0,772	0,73	0,868	0	0,868	0	0	1,063	0	0,333	0,333	0,835	0,333	0	

Longitude Bin	105	105	105	100	100	100	95	95	95	95	95	90	90	90	90	90	85	85	80	80	75	75	75	75	75	70	70	
Time of Day	D	N	N	N	D	Dk	D	Dk	N	D	N	N	D	D	N	N	Dk	N	N	D	D	N	D	Dk	N	Dn	N	
Number	500	503	504	506	508	509	511	512	513	514	515	516	517	518	519	520	523	524	527	529	531	532	534	535	536	537	539	
48	0	0	0	0	0	0	0	0	0,333	0	0	0	0	0	0	0	0,396	0	0,396	0	0	0	0	0	0	0	0	
49	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,333	0	0	0	
50	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
51	0,805	0	0,333	0,667	0	0,333	0	0	0	0	0,396	0	0	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	
52	0	0	0,333	0,793	0	1	0	0	0	0,396	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
53	1,294	1,294	1,063	1,168	0,667	0,333	0,73	2,459	0,868	1,524	1,39	0	0,73	1,376	0	0,333	0,333	0	0,333	0	1,264	0	0,439	0,439	0	1,201	1,231	
54	0,471	1,35	0	0,333	0	0,577	0	0,911	0	0	0	0,667	0,333	0,522	0	0	0	0	0	0	0	0,471	0	0	0	0	0	
55	0	0	0	0	0	0	0	0	0	0,439	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
56	1,543	1,29	0	0,772	0	0	0,333	0,793	0	0	0,333	0	0,333	0,471	0	0	0	0	0	0,333	0,333	0,333	0	0	0,333	0	0	
57	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,333	0	0	0	0	
58	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,439	0	0	
59	0,333	0	0	0,73	0	0,73	0,333	1,375	0	1,211	0,542	0,876	0	1,126	0	1,168	1,347	0	0,333	1,377	1,126	0,667	0	0,439	0,918	1,542	0,73	
60	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,396	
61	0,396	0,396	0	0,73	0	0,333	0	0,333	0	0	0,439	0	0	0	0	0	0	0	0	0	0	0	0	0,333	0	0	0	
62	0	1,201	0	1,824	0	1,214	0,333	0,832	0	0,667	0,894	0,805	1,379	0,396	0	1,105	0,91	0	0	0	0,333	1,612	0,333	0	1,59	1,377	0,97	
63	0	0,73	0	1,905	0	2,042	0	0,73	0,667	0,805	0,667	0,667	0	0,333	0,396	0,667	1,383	0,73	1,105	0	0,333	0,799	0,396	1,299	1,251	0	0	
64	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,396	0	0,498	0	0	
65	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,333	0	0	0	0	
66	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,333	0	0	0	0	
67	0	0	0	0	0	0,396	0	0	0	0	0	0	0,333	0,333	0	0	0	0	0	0	0	0	0,396	0	0	0	0	
68	0	0	0	0	0	0	0	0	0,333	0	0	0	0	0	0,333	0	0	0,333	0	0	0	0	0	0	0	0	0	
69	0	0	0	0	0	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0,667	0	0	0	0	0	0	
70	0	0	0	0	0	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,396	0	0	0,333	0	0
71	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,333	0	0
72	0,333	0,73	0	0,73	0	0,577	0	0	0	0	1	0,396	0,333	0	0,439	0,333	0	0,333	0,333	0	0	0,333	0	0,439	0,835	0	1,126	
73	0,396	0	0	0,333	0	0,835	0	0,333	0	0	0,333	0	0	0	0	0	0,396	0	0	0	0	0	0	0	0,333	0	0,667	
74	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
75	0	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,396	0	0	0	0,333	0	0	0,333	
76	0	0	0	0,333	0,73	0	0	0,835	0,333	0	0	0	0,333	0	0	0	0	0,333	0	0	0,439	0	0	0,333	0,835	2,031	0,396	
77	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
78	0	0,333	0	0	1,168	1,835	0,333	1,438	1,639	0,96	2,062	1,479	2,025	1,674	1,393	2,384	1,676	2,33	1,063	1,353	0,894	1,436	1,455	1,53	1,472	1,927	2,035	
79	0	0	0	0	0	0	0	1,063	0,73	0,333	0	0,333	0	0,333	0,333	0	0	0	0	0	0	0	0,333	0	0	0,333	0	0
80	1,642	3,701	0	4,44	0,73	3,655	0,333	1,517	2,917	0	2,731	2,477	2,38	0,333	2,166	1,168	2,173	0,333	2,215	0	0,439	1,245	0	2,336	0,73	0,396	4,28	
81	0	0	0	0	0,805	0	0	0	1,166	0	0	0,333	0	0	1,685	0	0	0,333	1	0	0	1,168	0	0,832	1,231	0	0	
82	2,624	3,925	0	6,203	0,333	2,957	0,333	0	2,412	0	1,3	2,196	1,946	0	1,88	1,439	0,793	0,439	1,396	0,471	0	0	0	1,217	1,059	0	1,71	
83	0,333	0	0	1,479	0,333	0,333	0	1,704	0	0,561	0,333	0	0,772	0,333	0	0	0	0,73	0	0,868	0,667	0	0,667	0,73	0,667	1,105	0,396	
84	0	0,656	0	0	0	0	0	0	0	0	0	0,333	0	0	0,333	0	0	0	0	0	0	0	0	0,333	0	0	0	
85	0	0	0	0	0	0	0	0,667	0	0	0,73	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
86	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,333	0	0	0	0	0	0,396	0	0,333	
87	0	0,333	0	0	0	0	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
88	0	0	0	0,577	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	1,244	0	0	
89	0	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,333	
90	0	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
91	0,522	0,471	0	0,522	0	0,396	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,333	
92	0	0	0	0	0	0,333	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
93	0	0,333	0	0,835	0	0	0	0	1,258	0	0,333	0	0	0	0,667	0	0,396	0	0	0	0	0	0	0	0	0	0,396	
94	0,498	0,772	1,063	1,314	0,396	0,868	0,667	0,333	1,347	0,333	0,805	0,73	1,168	0,333	1,333	0,73	0,868	1,238	0,471	0	1,063	0,793	0,333	0,868	0,835	1	1,274	

Longitude Bin	105	105	105	100	100	100	95	95	95	95	95	90	90	90	90	90	85	85	80	80	75	75	75	75	75	70	70	
Time of Day	D	N	N	N	D	Dk	D	Dk	N	D	N	N	D	D	N	N	Dk	N	N	D	D	N	D	Dk	N	Dn	N	
Number	500	503	504	506	508	509	511	512	513	514	515	516	517	518	519	520	523	524	527	529	531	532	534	535	536	537	539	
95	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,333	0	0	0	0	0	0	0	0	0	0,396	0	
96	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,333	0	0	0	0	0,667	0	0	
97	0	0,577	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
98	0	0	0	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
99	0	0,73	0	0,333	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
100	0	0	0	0	0	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
101	0	0	0	0	0	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
102	0	0	0	0	0	0	0	0,333	0,333	0	0,333	0	0	0,333	0	0	0	0	0,333	0,333	0	0	0	0	0	0	0	
103	1,138	0	0,835	1,534	2,516	1,856	1	3,065	1,82	2,472	2,114	1,73	2,087	2,013	3,857	1,805	2,646	3,286	2,445	1,77	2,132	2,843	2,256	2,915	2,72	2,031	2,015	
104	0,333	0	0,498	0	0	0	0	0	1,103	0,396	0	0	0	0,396	0,62	0	0	0	0,333	0	0	0	0	0	0	0	0,667	
105	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
106	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,333	0	0	0	0	0	0	
107	0	0,577	0	0,333	1,228	0	0	0	3,033	0,333	0	0,396	0,396	0	4,261	0,805	0,333	2,937	2,02	0	0,471	1,334	0	2,909	0,73	0,439	0	
108	0,73	0	0	0,96	0,439	0,333	1,063	1,032	0,876	0	0,522	0,73	0	0	0,868	0	0,439	0	0,793	0	0	0,333	0	0,667	0,396	0	0,667	
109	1,614	1,634	1,126	1,933	1,8	0,73	0,396	1,166	1,294	0,333	0,805	1,165	1,818	0	0,835	0,667	1,031	1,512	0,772	1,437	1,366	1,515	0,333	1,924	1,42	0,73	1,126	
110	0,333	0,333	0	0,73	0,396	0,333	0,333	0,333	0,333	0	0	0,333	0	0	0,333	0,333	0	0,73	0,333	0	0,333	0	0	1,063	0,73	0	0	
111	0,868	0	0	0,895	1,211	0,333	0,396	0,396	0,667	0,333	1,211	0	0,793	0,607	1,165	0	0	0,772	0	0	0,396	0,333	0	0,396	0,333	0	0,333	
112	0,73	0	0,471	0,396	0,832	0,333	0	0,333	0,333	0,877	0	0,333	1,272	0,333	0,667	0,396	0	0,577	0	0,805	1,287	0,561	0,957	0,522	1,833	0	0,333	
113	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
114	0,333	0	0	1,357	0,333	0,439	0	0	1,138	0	0	0,667	0	0	0,396	0,793	0	0,772	0	0	0	0	0	0,73	0,667	0	0	
115	0	0,772	0,333	0,333	0,835	0,396	0,667	1,51	1,748	1,078	1,02	0,333	1,412	1,253	2,33	1,27	0,805	1,745	0,667	0,772	1,201	1,703	0	1,498	1,696	0,772	1,231	
116	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,333	0	0	0,333	0	0,333	0	0,667	0,333	0	0	
117	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,333	0	0	0	0	0	0	0	0	0,333	0	0	0	
118	0	0	0	0	0	0	0	0	1,168	0,396	0	0	0	0	1,671	0	0	0,396	1	0	0	0	0	0,333	0	0	0	
119	0	0	0	0	0,396	0	0	0	1,347	0	0	0	0	0	0,926	0	0	0	0,73	0	0,333	0	0	0	0	0	0	
120	0	0,333	0,667	0,805	0,667	0,333	0	0	0,333	0	0	0	0	0,333	0	0	0	0	0	0	0	0	0,333	0	0	0	0,667	0
121	0,471	0,73	0	0	0	0,667	0	0	0,911	0	0	0	0	0,333	0	0,396	0	0,333	0,333	0	0	0	0	0,73	0,805	0	0,471	
122	0	0,439	0	0	0,396	0	0	0	0,73	0	0	0	0	0	0,667	0	0	0,333	0	0,396	0	0,333	0	0	0	0	0	
123	0	0	0	0	0	0	0	0	0	0,396	0	0	1,103	0	0	0	0	0,667	0	0	0	0	0	0	0,577	0	0	
124	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,333	0	0	
125	0	1,825	0,667	1,491	0	0,333	0,396	1,849	0,542	0,957	1,36	0,877	0	0,396	0,333	0	0,396	0	0	0	0	0	0	0,333	0,805	0,439	0,772	
126	0	0	0	0	0	0	0	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
127	0,333	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,73	
128	0	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	1,063	
129	0	0	0	0	0	0	0	0	0	0	0,333	0	0	0	0	0	0	0	0	0,333	0	0	0	0	0	0	0	
130	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0,333	
131	0	0	0	0	0	0,396	0	0	0	0	0	0	0	0	0	0,333	0	0	0	0	0	0	0	0	0	0	0,333	
132	0	0	1,063	1,091	0,333	1,382	0	0	0,333	0	0	0	0	0	0	0	0	0	0,333	0	0	0,333	0,333	0,333	0,73	0,333	1,168	
133	0,396	0,943	1,705	1,829	0,73	2,25	0,333	0	0	0	0	0	0,333	0	0,333	0,793	0	0,396	0,333	0	1,481	0,97	0,333	0,667	1,21	0,333	1,487	
134	0	0,868	1,127	0,498	0,989	1,254	0	0	0	0	0	0	0	0	0	0	0,333	0	0	0,333	0,333	0	0,439	0,396	0,805	0	1,031	
135	0,333	0,333	0,333	0,333	0,333	0,333	0	0	0	0	0	0	0	0,333	0,667	0	0	0	0,333	0	0	0,333	0	0,667	0,333	1,105	0	